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1 **High-pressure serpentinization and abiotic methanogenesis in metaperidotite from the**
2 **Appalachian subduction, northern Vermont**

3 Antoine Boutier^{1,2*}, Alberto Vitale Brovarone^{1,3,4}, Isabelle Martinez², Olivier Sissmann⁵, Sara
4 Mana⁶

5 ¹*Dipartimento di Scienze della Terra, Università degli Studi di Torino, Via Valperga Caluso 35,*
6 *Torino, Italy*

7 ²*Université de Paris, Institut de physique du globe de Paris, CNRS, F-75005 Paris, France*

8 ³*Dipartimento di Scienze Biologiche, Geologiche e Ambientali, Alma Mater Studiorum*
9 *Università di Bologna. Piazza di Porta San Donato 1, Bologna, Italy*

10 ⁴*Institut de Minéralogie, de Physique des Matériaux et de Cosmochimie (IMPMC), Sorbonne*
11 *Université, Muséum National d'Histoire Naturelle, UMR CNRS 7590, IRD UR206, 75005 Paris,*
12 *France*

13 ⁵*IFP Energies Nouvelles, 92500 Rueil-Malmaison, France*

14 ⁶*Department of Geological Sciences, Salem State University, Salem, MA 01970, USA*

15 *Corresponding author: antoine.boutier@unito.it

1 **Abstract**

2 Serpentinization is the process of hydroxylation of olivine-rich ultramafic rocks to produce
3 minerals such as serpentine, brucite and magnetite. This process is commonly accompanied
4 by Fe oxidation and release of H₂, which can be involved in abiotic reaction pathways leading
5 to the genesis of abiotic light hydrocarbons such as methane (CH₄). Examples of this
6 phenomenon exist at the seafloor, such as at the serpentinite-hosted Lost City hydrothermal
7 field, and on land in ophiolites at relatively shallow depths. However, the possibility for
8 serpentinization to occur at greater depths, especially in subduction zones, raises new
9 questions on the genesis of abiotic hydrocarbons at convergent margin and its impact on the
10 deep carbon cycle. High-pressure ultramafic bodies exhumed in metamorphic belts can
11 provide insights on the mechanisms of high-pressure serpentinization in subduction zones
12 and on the chemistry of the resulting fluids. This study focuses on the ultramafic Belvidere
13 Mountain complex belonging to the Appalachian belt of northern Vermont, USA.
14 Microstructures show overgrowth of both primary (Mg# 0.91) and metamorphic (Mg# 0.95)
15 olivine by delicate antigorite crystals, pointing to at least one stage of serpentinization at high-
16 temperature conditions and consistent with the high-pressure subduction evolution of the
17 Belvidere Mountain complex. Formation of ubiquitous magnetite and local Fe-Ni alloys
18 testifies to the partial oxidation of Fe²⁺ into Fe³⁺ and generation of reduced conditions. Fluid
19 inclusion trails cross-cutting the primary olivine relicts suggest their formation during the
20 antigorite serpentinization event. MicroRaman spectroscopy on the fluid inclusions reveals a
21 CH₄-rich gaseous composition, as well as N₂, NH₃ and H₂S. Moreover, the precipitation of
22 daughter minerals such as lizardite and brucite in the fluid inclusions indicate the initial
23 presence of H₂O in the fluid. High-pressure serpentinization driven by the infiltration of

24 metasediment-derived aqueous fluids is proposed at the origin of CH₄ and other reduced fluid
25 species preserved in the fluid inclusions. This suggests the Belvidere Mountain complex as an
26 example of deep abiotic hydrocarbon genesis related to high-pressure serpentinization in an
27 early Paleozoic subduction zone.

28 **Keywords:** HP Serpentinization, Metamorphic olivine, Abiotic methane, C fluxes at
29 subduction zones, Belvidere Mountain Complex.

1 Introduction

2 Alteration of ultramafic rocks by aqueous fluids, or serpentinization, produces serpentine-rich
3 rocks and may be accompanied by redox reactions leading to the formation of H₂ and abiotic
4 light hydrocarbons (Mével, 2003; Moody, 1976; Seyfried Jr et al., 2007). Oxidation of Fe²⁺ in
5 mantle minerals and formation of Fe³⁺ minerals such as magnetite produces H₂, which may
6 then promote the conversions of dissolved carbon-bearing species into abiotic CH₄ (Andreani
7 et al., 2013; Berndt et al., 1996). Serpentinization is well documented at mid-ocean ridges and
8 on land (Andréani et al., 2007; Cannat et al., 2010; Etiope et al., 2011; Klein et al., 2014;
9 Schrenk et al., 2013), where H₂ and CH₄ produced through serpentinization can support
10 biological communities and could be linked to the emergence of early life on Earth (Kelley et
11 al., 2005; Ménez et al., 2018; Sleep and Bird, 2007). Serpentinization processes may also
12 happen at much greater depths and affect subducted mantle sections and the overlying
13 mantle wedge at convergent margins (Guillot et al., 2015, 2000; Vitale Brovarone et al., 2017;
14 Wada et al., 2008), owing to the availability of aqueous fluids (Bebout and Penniston-Dorland,
15 2016; Deschamps et al., 2013). However, the patterns of serpentinization reactions at high-
16 temperature (HT) and high-pressure (HP) conditions, here defined with respect to the
17 serpentine stability field, and the associated fluid-rock redox budgets are still largely
18 unconstrained. The possibility that HT-HP serpentinization may not involve Fe oxidation and
19 genesis of H₂ has been proposed (Evans, 2010). However, natural case studies of slab-derived
20 serpentinized peridotites suggest that magnetite, H₂ and abiotic CH₄ can be produced at these
21 conditions (Vitale Brovarone et al., 2020). Ultramafic bodies in metamorphic belts can provide
22 good insights on the mechanisms of HP-HT serpentinization and the resulting fluids (Evans et

23 al., 2017), and therefore make good case studies for the investigation of the mechanisms of
24 serpentinization in subduction zones.

25 The Belvidere ultramafic complex, Vermont, USA, is a fragment of the Iapetus ocean that was
26 involved in the Taconic orogeny and recorded HP metamorphism during Cambrian-Ordovician
27 (Chew and van Staal, 2014; Gale, 1980; Honsberger et al., 2017). This complex includes
28 variably serpentinized peridotites bearing antigorite, the HT serpentine polymorph stable at
29 subduction zone HP conditions (Schwartz et al., 2013), therefore making a potential example
30 of HT-HP serpentinization. In this work, we integrate field, microstructural, thermodynamic,
31 and fluid inclusions data to investigate the patterns and timing of serpentinization of the
32 Belvidere ultramafic complex, with particular focus on the HP fluid-rock interactions recorded
33 by these rocks.

34 **1. Geologic setting**

35 The Belvidere Mountain complex (BMC) area is part of a North-South trending belt of
36 mafic/ultramafic rocks belonging to the Appalachian Mountain system and extending from
37 Newfoundland to Georgia (Gale, 2007; Hibbard et al., 2006)(Fig. 1). These ultramafic bodies
38 mainly consist of variably serpentinized peridotites, and associated metabasic and metafelsic
39 rocks tectonically embedded within Cambrian-Ordovician metasedimentary and
40 metavolcanic formations (Chidester et al., 1978; Gale, 1986, 1980; Laird et al., 1984; Van
41 Baalen et al., 2009). The BMC, as well as other mafic/ultramafic complexes such as the
42 Tillotson Peak and the Pennington complexes, are interpreted to represent remnants of an
43 ocean-continent transition zone associated with the extension of the Laurentian margin
44 (Chew and van Staal, 2014). These terranes were successively involved in the closure of the

45 Iapetus Ocean during the Ordovician Taconic orogeny (Doolan et al., 1982; Honsberger et al.,
46 2017; Karabinos et al., 1998; Laird et al., 1984; Stanley et al., 1984). The BMC forms a 6.5 km
47 long body that structurally overlies the metasedimentary Ottawaquechee and Stowe
48 Formations to the East and is overlain by the Hazen Notch formation to the West. The Stowe
49 Formation is late Cambrian, composed of gray-green quartz-chlorite-sericite +/- magnetite
50 schists. The Ottawaquechee Formation is middle Cambrian consisting carbonaceous
51 pyritiferous phyllite. The Hazen Notch Formation is dated from Neoproterozoic to Cambrian,
52 graphitic and non-graphitic quartz-albite-sclerite-chlorite schist, along medium-grained,
53 massive, quartz-albite-muscovite gneiss. The internal architecture of the BMC comprises,
54 from top to bottom, variably serpentinized dunite and harzburgite, coarse-grained and fine-
55 grained amphibolite, greenstone, muscovite schist, and albite gneiss (Gale, 1986, 1980; Van
56 Baalen et al., 2009). Lenticular bodies composed of talc-carbonate rocks and steatite, often
57 including cores of relatively un-serpentinized peridotite, are observed within the BMC, and
58 range size from a few meters to several tens of meter (Chidester et al., 1978). The BMC rocks
59 underwent Taconic blueschist-facies peak metamorphic conditions constrained at 0.9-1.3 GPa
60 and 510-520 °C in metabasic rocks (Honsberger, 2015; Laird et al., 1993). Slightly higher P,
61 blueschist-to-eclogite-facies metamorphic conditions are recorded by the Tilliston Peak mafic
62 rocks located just north of the BMC (Laird et al., 1993, 1984). This tectonometamorphic event
63 has been dated at 505-473 Ma by ^{40/39}Ar amphibole and mica geochronology (Castonguay et
64 al., 2012; Laird et al., 1993).

65 **2. Methods**

66 Quantitative energy-dispersive X-ray spectroscopy (EDS) analyses and backscattered electron
67 (BSE) imaging were carried out using the JSM-IT300LV Scanning Electron Microscope Oxford

68 Inca Energy Dispersive Spectrometer at the Department of Earth Sciences of Turin University.
69 Data were processed with the INCA software from Oxford Instruments. Quantitative analysis
70 employed 15kV accelerating voltage and 20 s to 40 s counting time. Natural and synthetic
71 mineral and oxide standards were employed. EDS calibration was made using Cobalt
72 standard.

73 Quantitative wavelength-dispersive spectrometer (WDS) analyses were carried out using a
74 JEOL 8200 Super Probe at the Department of Earth Sciences "Ardito Desio" of Milan
75 University. The microprobe was using a 15keV accelerating voltage under 5 nÅ, with 30s
76 counting time under maximum emission peak. Sixteen oxide composition were measured,
77 using synthetic and natural standards: grossular (Si, Al and Ca), omphacite (Na), K-feldspath
78 (K), fayalite (Fe), forsterite (Mg), rhodonites (Mn), niccolite (Ni), ilmenite (Ti), galena (Pb and
79 S), pure Cr, pure Zn and pure Cu.

80 MicroRaman spectroscopy of minerals and fluid inclusions was done at the Department of
81 Earth Sciences, University of Turin, with a LabRAM HR (VIS) (HORIBA Jobin Yvon) equipped of
82 a 532.11 nm, solid-state Nd laser, a Super Notch Plus filter with spectral resolution of 1 cm⁻¹,
83 and a grating of 600 grooves/mm. The laser of emission power was set at 80 mW and focused
84 to 5 μm with a ×100 objective with a laser power on the sample < 5 mW. Calibration was
85 performed using the 520.6 cm⁻¹ band of a silicon standard for the 100-2000 cm⁻¹ range, and
86 the 2331 cm⁻¹ band of atmospheric N₂ for the 2000-4000 cm⁻¹ range. Four accumulations of
87 30–60 s were collected for each spectrum. Raman spectra of fluid inclusions were performed
88 on double-polished thick sections unless otherwise specified.

89 Isotopic composition measurements of methane were performed on a MAT 253 (Thermo
90 Fisher) mass spectrometer, coupled with gas chromatography (GCC-IR-MS) in order to purify
91 the samples. The gas phase was extracted and measured after crushing the rock sampler
92 under vacuum.

93 Thermodynamic modelling was performed using the Deep Earth Water (DEW) Model
94 (Sverjensky et al., 2014) and the EQ3/EQ6 software (Wolery and Jarek, 2003) with a modified
95 Berman database (Berman, 1988). EQ3 was used to calculate the composition of a fluid in
96 equilibrium with a given mineral assemblage as, fO_2 , P and T. EQ6 was used to model the
97 interaction of fluid compositions from EQ3 with another mineral assemblage as function of
98 the fluid/rock (F/R) ratio.

99 **3. Sample description and mineral chemistry**

100 In this section, we present the petrography, microstructures and mineral chemistry of
101 samples collected in the Belvidere mine (Fig. 2) and showing different degrees of
102 serpentinization. The selected samples range from weakly serpentinized dunite (V18-2a and
103 V18-2b) to partially serpentinized dunite including layers of boudinaged meta-pyroxenite
104 (V18-3a and V18-3b), to fully serpentinized peridotite (V18-B3).

105 **3.1. *Microstructural characterization***

106 Sample V18-2a (least serpentinized dunite) consists of olivine (~80 vol.%), antigorite (~10
107 vol.%), Cr-spinel (>5 vol.%), magnetite (<5 vol.%), chlorite (<1 vol.%), sulphides (<1 vol.%), and
108 alloys (<1 vol.%) determined from visual estimate in thin section. Primary olivine forms
109 crystals ranging in size from 0.25 to 1 mm (Fig. 3A). It is partially replaced by antigorite, as

110 identified by MicroRaman spectroscopy (Fig. 4), along mesh-like structures (Fig. 3A-C).
111 Antigorite is present as elongated crystals, up to 200 μm in length, which statistically
112 overgrew the olivine (Fig. 3C). Magnetite is present in three microstructural domains: as large,
113 millimeter scale crystals rimming chromite relicts, as grains of $\sim 50 \mu\text{m}$ in the mesh structures,
114 and as millimeter scale crystals in antigorite veins. Chlorite is found, together with magnetite,
115 around Cr-spinel relicts (Fig. 3B).

116 Sample V18-2b (serpentinized dunite) consists of antigorite ($\sim 40 \text{ vol.}\%$), olivine ($\sim 40 \text{ vol.}\%$,
117 including primary and metamorphic olivine), magnetite ($>5 \text{ vol.}\%$), chlorite ($>5 \text{ vol.}\%$), brucite
118 ($<5 \text{ vol.}\%$), and alloys ($<1 \text{ vol.}\%$). The structure is similar to V18-2a, but in this case the extent
119 of serpentinization along the meshes is higher. SEM backscattered-electron imaging reveals
120 the presence of a second generation of olivine, hereafter metamorphic olivine (see Section
121 4.2 for discussion) growing in two different sites: epitaxially on the primary olivine and
122 replacing former orthopyroxene (Fig. 5A-B). Striped zoning is observed in primary olivine (Fig
123 5A), as described in Plümper et al., 2012a, alternating thin forsterite-richer and forsterite-
124 poorer olivine composition. The initial presence of orthopyroxene in the rock is inferred based
125 on the presence of clinopyroxene exsolutions preserved in metamorphic olivine-rich
126 pseudomorphoses (Fig. 3D and 5B). Antigorite is present as elongated (0.5 mm in length)
127 crystals and as fine-grained aggregates in the matrix. The elongated antigorite crystals are
128 chemically zoned, with bright core and dark rim in backscattered electron imaging (Fig. 5A).
129 Raman spectra of both generations exhibit the characteristic 1043 cm^{-1} band of antigorite,
130 whereas the main OH stretching is at 3664 cm^{-1} in the bright core and at 3673 cm^{-1} in the dark
131 rim (Fig 4C), both inconsistent with antigorite-lizardite mixing. The higher Raman shift of the
132 dark rim may be linked to variation of pressure condition (Auzende et al., 2004). The dark

133 antigorite generation appears the same forming the fine-grained aggregates. The bright
134 antigorite generation (at the core of large crystals) is never found in contact with either
135 primary or metamorphic olivine, whereas the darker antigorite generation (at the rim of large
136 crystals of disseminated in the matrix) is in contact with them (Fig. 5A). Magnetite is more
137 abundant relative to sample V18-2a, especially along meshes. Brucite was identified by SEM
138 and Raman (Fig. 4D) in veins, sometimes associated with metamorphic olivine (Fig. 5C). Alloys
139 and sulphides are present in association with antigorite and are locally associated to
140 magnetite (Fig. 5D).

141 Samples V18-3a and 3b are from a boudinaged layer of clinopyroxenite included in an
142 intensely serpentinized dunite (Fig. 3E-G). The mineral assemblage and microstructures of the
143 serpentinized dunite part of the sample is similar to sample V18-2b, yet more intensely
144 serpentinized. The primary clinopyroxene is fully replaced by diopside aggregates in both the
145 clinopyroxenite layer and the host dunite. The primary clinopyroxene sites in this sample are
146 slightly different from the serpentinized dunite, with less abundant magnetite and
147 characteristic fan-shaped diopside aggregates in the former compared to the latter.
148 Metamorphic olivine is present at the rim of primary clinopyroxene and along its cleavages,
149 together with antigorite (Fig. 5E-F). In the latter case, metamorphic olivine is localized along
150 thin arrays encircled by antigorite (Fig. 5F). Antigorite shows the same chemical zoning as
151 observed in the dunite, with brighter cores and darker rims in backscatter electron imaging
152 (Fig. 5G). Backscattered electron imaging reveals that the dark antigorite shows similar
153 overgrowth microstructural relationships with both primary and metamorphic olivine (Fig.
154 5E). The microstructures reflect either equilibrium between antigorite and the two olivine
155 generations, or overgrowth of both olivine generations by the antigorite. Considering that the

156 amount of primary olivine strongly decreases in favour of antigorite in several samples, which
157 suggests serpentinization of the primary olivine, second antigorite generation formed after
158 the growth of metamorphic olivine appears . It is possible, even though the microstructures
159 could not confirm it, that the first antigorite generation grew prior to or together with the
160 metamorphic olivine. Magnetite is present in four different structural sites: as arrays of ~50
161 μm size crystals along the meshes, as trails of ~50 μm crystals in primary olivine and
162 clinopyroxene pseudomorphoses, as aggregates rimming the primary Cr-spinel, and as
163 millimeter scale aggregates along discordant veins. The Cr-spinel sites are characterized by
164 three different layers: a rather preserved core, a mantle of ferritchromite, and a rim of
165 magnetite (Fig. 5E; see Section 4.2).

166 Sample V18-B3 is a fully serpentinized peridotite. The matrix is composed of a mixture of
167 antigorite and chrysotile (identified by MicroRaman) with magnetite. Brucite is observed
168 replacing pyroxene sites and is associated with magnetite and/or metamorphic olivine (Fig.
169 3H and 5H). Based on the microstructural observations on sample V18-2b, these
170 pseudomorphoses are interpreted as former orthopyroxene crystals replaced by
171 metamorphic olivine and successively hydrated to form serpentine + brucite \pm magnetite.

172 **3.2. Mineral chemistry**

173 Primary olivine has Mg# of 0.91-0.92 [Mg# = $\text{Mg}/(\text{Fe}+\text{Mg})$], whereas metamorphic olivine is
174 enriched in Mg (Mg# of 0.95) (Table 1, Fig. 6). The Mn# [Mn# = $\text{Mn}/(\text{Mn}+\text{Fe}+\text{Mn}+\text{Ni})$] of
175 metamorphic olivine (0.0040 ± 0.0006) is much higher than primary olivine (0.0015 ± 0.0005).
176 Metamorphic olivine analyses in sample V18-B3 show slightly lower Mg# and higher Mn#
177 content relative to metamorphic olivine from other samples. Metamorphic olivine rimming

178 the primary orthopyroxene sites in sample V18-3a has higher CaO content (0.25 wt.%)
179 compared to the metamorphic olivine overgrowing primary olivine in V18-2b (0.02 wt.%).
180 Metamorphic olivine in the sample V18-B3 show slightly lower Mg# and highly increased MnO
181 content (1.2 wt %) in regard to metamorphic olivine from other samples. The NiO content of
182 primary olivine (0.38 to 0.49 wt.%) and metamorphic olivine (0.39 to 0.46 wt.%) are similar.

183 The incorporation of Mn in olivine appears to be characteristic of metamorphic olivine from
184 several localities regardless of the olivine formation environments (dehydration vs.
185 hydration). The Mg# of metamorphic olivine relative to primary olivine may depend upon
186 several parameters such as the stability of different Fe-bearing minerals, the Fe partitioning
187 among them, and the P-T and redox conditions (Frost and Beard, 2007; Majumdar et al., 2016;
188 Nozaka, 2018, 2003). An increase in Mg# (and MnO) in metamorphic olivine relative to
189 primary olivine has been observed in inferred mantle wedge peridotites recording HP
190 serpentinization (Dandar et al., 2019; Guillot et al., 2000). Plümper et al., 2012a report striped
191 Mg# zoning in hydrated supra-subduction mantle rocks as the result of chemical interaction
192 during antigorite serpentinization at high temperature conditions. Metamorphic olivine
193 formed through prograde antigorite + brucite dehydration shows either lower or higher Mg#
194 compared to mantle olivine (Arai et al., 2012; Debret et al., 2013; Iyer et al., 2008; Kempf and
195 Hermann, 2018; Nozaka, 2018; Plümper et al., 2012b; Scambelluri et al., 1995; Shen et al.,
196 2015)(Fig. 6).

197 The core of individual, elongated antigorite crystals exhibits higher FeO (1.6 to 2.6 wt.% [Mg#
198 0.94-0.96]), Al₂O₃ (1.77 wt.%) [Mg# =Mg/(Mg+ΣFe)], and Cr₂O₃ (0.61 wt.%) than the rim (1.35
199 wt.% [Mg# 0.97], 0.52 wt.%, 0.13 wt.%, respectively) (Table 1). Unzoned matrix antigorite has

200 a composition equivalent to the rim of the elongated antigorite crystals. The Mn# [Mn# =
201 Mn/(Mn+Fe+Mn+Ni)] of antigorite is 0.0006 (± 0.0004).

202 The preserved spinel core has Cr# [Cr/(Cr+Al)] of 0.79 (Table 2). The mantle overgrowing the
203 primary spinel has a ferritchromite mantle of Cr-magnetite with a Cr# of 0.91 and higher MnO
204 relative to the core. The magnetite rims with no Al₂O₃, but high Cr₂O₃ (1.87 wt.%). The
205 composition of magnetite in the meshes and veins differs from magnetite found in the spinel
206 sites, with very little Cr₂O₃.

207 Diopside has a Mg number [Mg# =Mg/(Mg+Fe)] of 0.98. Chlorite has an Mg# [Mg#
208 =Mg/(Mg+ΣFe)] of 0.95, and Cr₂O₃ up to 4.07 wt.% (Table 1).

209 Brucite has up to 3 wt.% FeO (as total Fe), Mg# [Mg# =Mg/(Mg+Fe)] of 0.97 (Table 2). Alloys
210 and sulfides are Ni rich, with various amounts of S, Fe and trace amounts of Pb, Cu, et Co
211 (Table 3). Alloys mainly consist of Cu-bearing FeNi alloy (taenite) (Table 3). Sulphides include
212 Ni, Cu, and Fe sulphides. Only heazlewoodite (NiS) was analysed whereas for other sulphides
213 proper analyses could not be obtain owing to the small grain sizes.

214 **3.3. Fluid inclusion analysis**

215 Primary olivine in all samples is rich in fluid inclusions forming secondary trails (Fig. 7). The
216 fluid inclusion trails are confined within individual crystals and show two alternative structural
217 relationships relative to the antigorite veins, being either cut by them (Fig. 7A) or injected
218 from them (Fig. 7B-C). These patterns suggest a secondary fluid trapping during the antigorite
219 serpentinization event forming the main mesh structure observed in the rock.

220 Fluid inclusions exhibit rounded to elongated shapes (Fig. 7D-E). Optical microscope
221 observations at room conditions suggest that the fluid inclusions are either single-phase and
222 gaseous, or bi-phase with solid and gas. MicroRaman spectra of fluid inclusions are presented
223 in Fig. 8. The spectra show the presence of marked CH₄ bands (2912 cm⁻¹), as well as N₂ (2327
224 cm⁻¹), NH₃ (3324 cm⁻¹), S-H/H₂S (2575 cm⁻¹). Free H₂O in the fluid inclusions was not detected
225 by MicroRaman. However, the presence of tiny amounts of free H₂O in the fluid inclusions —
226 undetectable by MicroRaman at room conditions— cannot be excluded (Berkesi et al., 2009;
227 Lamadrid et al., 2017).

228 The molar fraction of gas in the gas mixture was estimated using the Raman scattering cross-
229 section and the instrumental efficiency of each species. Using equation presented in Frezzotti
230 *et al.*, 2012, we obtain the following molar proportion: CH₄ = 92±6%, N₂ = 6±5%, H₂S = 1±1%
231 and NH₃ = 1±2% for the mean composition of twenty inclusions in olivine from all samples (Fig
232 9). Composition within the same inclusion trail show molar fractionation variation smaller
233 than 1% but different trail within the same crystal can exhibit up to 10% of molar fraction of
234 CH₄ and could reflect fluid heterogeneity. Alternatively, this effect can be the result of
235 different crystal orientation during data acquisition (Caumon et al., 2019). No clear pattern
236 was observed linked to relationship between inclusion trails and antigorite veinlets.

237 Solid phases were identified in large fluid inclusions. MicroRaman analysis revealed the
238 presence of lizardite and brucite but no magnetite was detected. Some inclusions contain
239 graphite as a solid phase in the inclusion in addition (Fig. 7F and Fig. 8B). The only sporadic
240 presence of graphite in the fluid inclusions, and in particular in inclusions resulting from
241 necking processes, suggest that this mineral precipitated as a result of local respeciation of
242 the fluid inclusions (Cesare, 1995). The presence of hydrous phases in the inclusions suggests

243 re-equilibration of the inclusion with the host mineral, effectively serpentinizing the olivine
244 and consuming water initially present in the inclusions. However, the presence of step-
245 daughter minerals in the inclusion is not systematic, suggesting that the initial fluid was
246 already rich in CH₄ (-N₂-NH₃-H₂S) at the time of trapping, and that the reduced fluid species
247 did not form only inside the fluid inclusions.

248 **3.4. Rodingite**

249 Rodingite are ultramafic rocks that are composed of carbonate, garnet, diopside, epidote ±
250 graphite. While not being the focus of this study, the carbonate in the rodingite bear
251 numerous trails of fluid inclusions. Fluid inclusions exhibit negative crystal shape and are
252 single-phase gaseous. The composition of the carbonate-hosted fluid inclusions, as revealed
253 by Raman spectroscopy, is essentially CH₄ and N₂. Using quantitative estimation of inclusions
254 in carbonates (see section 4.3) yield molar fractions of N₂ = 67±4%, CH₄ = 33±4% on four
255 different inclusions.

256 **3.5. Thermodynamic modelling**

257 Thermodynamic calculations were performed in order to constrain the mineralogical, fluid
258 and redox pattern of the HP serpentinization. Because the selected partially serpentinized
259 samples are comprised in strongly serpentinized rocks and embedded in metasediment, two
260 fluid equilibriums can be considered. For silica-rich fluid sources (e.g. metasedimentary rocks
261 or talc-bearing ultramafic rocks), the predicted assemblages do not match the natural
262 samples. For this reason, the fluid composition calculated with EQ3 was equilibrated with a
263 serpentinite consisting of antigorite + magnetite + brucite + chlorite + olivine. This fluid then
264 reacted with a dunite with mineral modal proportions comparable with the study samples

265 and consisting of 92% olivine (forsterite 90%, fayalite 10%), 5.4% orthopyroxene (enstatite
266 90%, ferrosilite 10%), 2.5% clinopyroxene (diopside 90%, hedenbergite 10%) and 0.1% spinel.
267 The fO_2 of the infiltrating fluid was set at the quartz-fayalite-magnetite (QFM) buffer. The
268 calculations were done at temperatures consistent with the peak metamorphic conditions
269 estimated for the BMC, i.e., 400 °C to 500 °C and 1 GPa (Honsberger, 2015; Laird et al., 1993),
270 and for different F/R ratios. Figure 10 shows the mineralogical evolution as a function of
271 reaction progress at 450 °C and 1GPa for a fluid rock ratio of 1. The model reaction proceeds
272 with progressive transformation of, from the first to the last reacting mineral, mantle spinel,
273 clinopyroxene, orthopyroxene, and olivine. Magnetite starts to form during the early stages
274 of the reaction along with chlorite in response to spinel consumption. Reaction of mantle
275 pyroxenes marks the precipitation of metamorphic clinopyroxene (diopside 93%,
276 hedenbergite 3.35% and clino-enstatite 3.65%), metamorphic olivine, antigorite, and
277 additional magnetite. The late formation of antigorite marks the partial consumption of
278 metamorphic olivine and a decrease in its Mg#, from Mg# 0.90 to Mg# 0.83. These patterns
279 reflect the microstructural features observed in the natural samples. For example, spinel
280 appears intensely replaced by chlorite in rather unserpentinized portions of the rock (Fig.
281 10A). Similarly, mantle olivine adjacent to fully replaced mantle pyroxenes is commonly little
282 affected by the serpentinization. Both mantle and metamorphic olivine in the natural samples
283 appear texturally replaced by antigorite, as also suggested by the modelling while
284 approaching equilibration. With the chosen bulk composition and used thermodynamic data
285 set, the formation of antigorite is limited to $T < 470$ °C. The antigorite stability field could be
286 enlarged by considering iron in the antigorite thermodynamic model. The fO_2 decreases
287 progressively during the reaction progress, with a steep decrease during the formation of
288 antigorite down to $\Delta\text{Log QFM} = -1.2$ (Fig. 10C). Methane, initially about 1 order of magnitude

289 less concentrated than CO₂ in the reacting fluid, becomes a dominant species at reaction
290 completion (about 1 order of magnitude more concentrated than CO₂). Nevertheless,
291 variations of parameters such as F/R ratio, mineralogy, or temperature, were found to affect
292 the proportion of CH₄ and CO₂. As an example, increasing the modal proportion of either
293 orthopyroxene or clinopyroxene in the initial rock or decreasing the F/R ratio appears to
294 favour a higher proportion of CH₄ in the fluid in respect to CO₂. The model predicts high
295 relative concentration of H₂ in the fluids, however H₂ was not detected in natural samples by
296 micro-Raman spectroscopy. Sulphur and nitrogen speciation were also assessed, with HS⁻ and
297 H₂S and NH₃ being the dominant sulphur species relative to HS O₄⁻ and N₂ and NH₄⁺,
298 respectively (Fig. 10B).

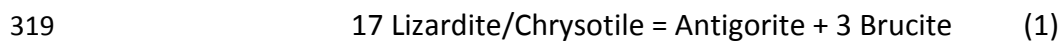
299 **4. Discussion**

300 **4.1. Patterns and timing of serpentinization**

301 The timing of serpentinization of slab-derived, exhumed HP serpentinized rocks can span
302 (sub)seafloor conditions prior to subduction, prograde hydration during subduction, or
303 retrograde hydration during exhumation. The BMC complex has been interpreted as a
304 fragment of subducted Iapetus lithosphere, and therefore may have recorded different stages
305 of serpentinization. Although the possibility of at least some (sub)seafloor serpentinization
306 prior to the Taconic subduction cannot be excluded, our data cannot provide any proof of
307 such a pre-subduction event and, instead, suggest a main hydration event at HP-HT conditions
308 in the subduction zone.

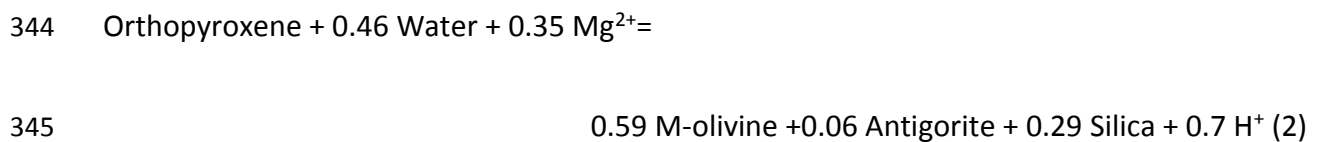
309 Besides late chrysotile veinlets, antigorite is the only serpentine polysome identified in the
310 BMC rocks. Although formation of antigorite may occur in a wide range of P-T conditions also
311 as a function of chemical parameters such as the silica activity (Rouméjon et al., 2019),

312 serpentinites dominated by antigorite are generally referred to the HT temperature part of
313 the serpentine stability field, generally above ~300-400 °C (Evans, 2004; Schwartz et al.,
314 2013). In most subduction zone settings, these conditions also correspond to relatively HP
315 conditions above 1 GPa. However, the presence of antigorite does not necessarily imply the
316 serpentinization event to have happened at HP-HT conditions because it could also have
317 formed as a result of the prograde transformation of lizardite or chrysotile following the
318 reaction:

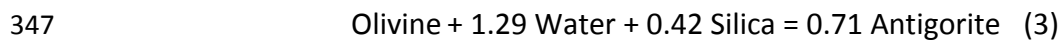


320 Nevertheless, several lines of evidence indicate that the BMC rocks recorded a stage of HP-
321 HT serpentinization. As a first general consideration, the presence of fresh mantle
322 assemblages throughout the complex (Chidester *et al.*, 1978; this study) represents a suitable
323 condition to promote HP-HT serpentinization in the subduction zone, if aqueous fluids are
324 available. Such a process has already been proposed in other HP ultramafic massifs preserving
325 fresh mantle assemblages (e.g. Früh-Green et al., 2004; Scambelluri and Tonarini, 2012; Vitale
326 Brovarone et al., 2020, 2017). The BMC rocks provides evidence for such a HP hydration event.
327 Figure 11 summarizes the proposed fluid-rock evolution of the BMC rocks as inferred from
328 the studied samples. The first indication of serpentinization (stage I in Fig. 11) is suggested by
329 the growth of elongated metamorphic olivine on the primary clinopyroxene sites (Fig. 5G),
330 which suggests the former presence of serpentine blades overgrowing the primary
331 clinopyroxene. The selective growth of serpentine at the expense of clinopyroxene rather
332 than primary olivine places this hypothetical event at HT conditions (Klein et al., 2013) and
333 possibly in the antigorite stability field (Fig. 12). Metamorphic olivine is most commonly
334 interpreted to form in response of serpentine dehydration during prograde metamorphism

335 (Plümper et al., 2017; Scambelluri et al., 1991). Figure 12 provides a compilation of traditional
336 serpentine dehydration reactions leading to the formation of metamorphic olivine. The peak
337 metamorphic conditions proposed for the BMC (0.9-1.3 GPa and 520 °C) are consistent with
338 the first olivine-forming reaction involving antigorite + brucite as reactants. Moreover, several
339 studies have shown that olivine can form at T conditions lower than the reactions shown in
340 figure 12 as a result of local bulk compositional features (Plümper et al., 2017). Alternatively,
341 metamorphic olivine after orthopyroxene may also have formed in response to hydration
342 rather than dehydration reactions, as already proposed for serpentized mantle wedge rocks
343 (Dandar et al., 2019) following the reaction:



346 Followed by the hydration of olivine to form antigorite following the reaction



348 In either case (hydration or dehydration), the amount of serpentization predating the
349 formation of metamorphic olivine must have been very low and, if any, related to a HT event.

350 The second, more robustly constrained stage of transformation (stage II in Fig. 11) is
351 characterized by the formation of metamorphic olivine after primary orthopyroxene, and as
352 rim around primary clinopyroxene (now recrystallized into diopside). This reaction was
353 observed in rocks containing rather undisturbed primary olivine, which again point to HT
354 serpentization conditions (Klein et al., 2013). The partial preservation of clinopyroxene
355 exsolution lamellae inside metamorphic olivine pseudomorphic on orthopyroxene may
356 suggest nearly isovolumetric replacement during this event (Plümper et al., 2012b; Viti et al.,

2005). The thermodynamic modeling results suggest that, at 450 °C and 1 GPa, a transient antigorite generation may have formed together with metamorphic olivine early in the fluid-rock interaction, along with spinel breakdown. The Al-Cr-rich antigorite cores (Atg1) observed in the studied samples may testify to this transient antigorite formation. The application of empirical distribution coefficient K_D for antigorite and olivine by Evans, 2008 and Trommsdorff and Evans, 1974 support this hypothesis (Table 4). The empirical Mg distribution coefficient K_D [antigorite/olivine, $K_D = (\sum Fe/Mg_{Srp})/(Fe/Mg_{Ol})$] proposed by Evans, 2008 is 0.45-0.35 upon consideration of ferric iron in the antigorite. The Atg1/metamorphic olivine apparent K_D is 0.49-0.25, which agrees with the value proposed by Evans, 2008 for equilibrium, compared with other antigorite/olivine pairs (Atg1/primary olivine = 0.14-0.25; matrix antigorite (Atg2)/primary olivine = 0.11; Atg2/metamorphic olivine = 0.17-0.25). The Mn distribution between antigorite and olivine gives similar results. The empirical Mn distribution coefficients for antigorite/olivine equilibria is 0.18 (Trommsdorff and Evans, 1974). The Atg1/metamorphic olivine pair has apparent K_D of 0.14-0.18, thus consistent with the predicted equilibrium (Table 4).

The third stage (stage III in Fig. 11) is characterized by the growth of a matrix antigorite (Atg2) at the expense of both primary and metamorphic olivine. This event may mark either retrograde hydration along the exhumation path of the BMC, or the progression of the fluid-rock interaction. As indicated by the thermodynamic modeling results, a second antigorite generation is expected to form at 450 °C and 1 GPa after the precipitation of metamorphic olivine (Fig. 10A). The modeling predicts this second antigorite and metamorphic olivine to be stable together. However, the apparent K_D for matrix antigorite and metamorphic olivine suggest disequilibrium (equilibrium Mg $K_D = 0.45$, measured = 0.17-0.25; equilibrium Mn $K_D =$

380 0.18, measured = 0.09-0.14) (Table 4). This feature suggests that stage III marks the beginning
381 of retrograde hydration.

382 The fourth stage of serpentinization (IV in Fig. 11) is characterized by the complete
383 serpentinization of primary olivine and partial to full serpentinization of metamorphic olivine.
384 The partial preservation of metamorphic olivine after orthopyroxene in sample free of any
385 primary olivine relict indicates that the largest event of serpentinization of the BMC rocks
386 took place after the formation of metamorphic olivine. The presence of both antigorite and
387 lizardite + brucite at the expense of metamorphic olivine after orthopyroxene suggests that
388 this event protracted during the cooling of the BMC metamorphic path to temperatures lower
389 than 400 °C (Fig. 12).

390 In summary, based on the collected petrographic and thermodynamic data, the dominant
391 serpentinization event observed in the BMC complex is interpreted to have taken place in the
392 Taconic subduction zone.

393

394 **4.2. Timing of fluid inclusion formation and origin of CH₄**

395 **4.2.1. Timing of fluid inclusion entrapment**

396 Fluid inclusions in the BMC ultramafic and related metasomatic rocks contain reduced fluid
397 species such as CH₄, NH₃, and H₂S. The timing of fluid inclusion formation, as well as the origin
398 of their reduced speciation, is discussed in this section.

399 Reduced fluids have been found to form in ultramafic systems in a wide range of geologic
400 conditions spanning mid-ocean ridges, obducted ophiolitic massifs, and orogenic peridotite
401 bodies, in subduction, and in the upper mantle (Andréani et al., 2007; Etiope et al., 2011;

402 Schrenk et al., 2013). In the BMC rocks, the fluid inclusions may have formed in three different
403 stages of the evolution of the massif and corresponding to three different geodynamic
404 settings: (i) in the mantle prior to the formation of the Iapetus Ocean, (ii) during the
405 (sub)seafloor evolution prior to subduction, and (iii) in the subduction zone.

406 A primary mantle origin can be ruled out because the observed fluid inclusions occur as
407 secondary trails propagating from the antigorite veinlets crossing olivine crystal. A
408 (sub)seafloor origin would match the identification of CH₄-rich fluid inclusions in oceanic
409 peridotite (Cannat et al., 2010; Holm and Charlou, 2001). However, in this case, the fluid
410 inclusions would have been preserved during prograde metamorphism to at least 520 °C and
411 1 GPa, which is unlikely (Touret, 2001). Moreover, the BMC fluid inclusions are rich in N
412 species (NH₃, N₂), which seems to be an uncommon feature in fluid inclusions from oceanic
413 peridotites (Grozeva et al., 2020; Klein et al., 2019). Instead, N-rich, NH₃-bearing fluid
414 inclusions have been recently documented in CH₄-rich fluid inclusions in HP serpentized
415 peridotites from the Alpine belt and proposed to represent a distinctive feature of subduction
416 zone serpentizing fluids relative to mantle or (sub)seafloor fluids, especially in the presence
417 of metasediment-derived fluids (Vitale Brovarone et al., 2020). Considering the
418 microstructural and petrologic features discussed in Section 5.1, and the abundance of N
419 species, a metamorphic origin in the Taconic subduction is proposed for the studied fluid
420 inclusions.

421 The formation of lizardite and brucite as step-daughter minerals inside the fluid inclusions can
422 be interpreted (1) as a prograde, pre-antigorite inclusion-host interactions, (2) as a retrograde
423 reequilibration of antigorite during cooling of the BMC below ~400 °C with excess brucite
424 (Reaction 1), or (3) as a retrograde host-inclusion interaction below ~400 °C (Fig. 12).

425 Excluding lizardite (meta)stability due to local equilibrium/kinetic features or faster
426 antigorite-lizardite conversion in the inclusions relative to the host rock, we interpret the
427 formation of step-daughter lizardite and brucite as a retrograde host-inclusion interaction.
428 Similar interpretations have been proposed for analogous inclusions from Alpine belt (Vitale
429 Brovarone et al., 2020).

430

431 **4.2.2. Origin of the reducing potential and fluid sources**

432 Another important question is the origin the identified reduced fluids species. In particular,
433 several studies over the last decades have investigated the biotic or abiotic origin of geological
434 CH₄ and associated reduced fluid species (Etiope et al., 2011; Etiope and Sherwood Lollar,
435 2013; McCollom, 2016; Ménez et al., 2018). Ultramafic systems are generally favourable
436 environments for the genesis of abiotic CH₄ (Section 5.2.1), but other interpretations are also
437 possible.

438 For example, the abundance of metasedimentary rocks in the study area may have promoted
439 the formation of thermogenic gases during their prograde evolution. This hypothesis would
440 be also consistent with the production of NH₃ through degassing of organic matter in
441 metasedimentary rocks (Bebout and Fogel, 1992; Li et al., 2009). Biotic processes, including
442 thermogenic gas formation, typically show very light $\delta^{13}\text{C}_{\text{CH}_4}$ signatures (Etiope and Sherwood
443 Lollar, 2013), whereas abiotic processes generally result in much heavier $\delta^{13}\text{C}_{\text{CH}_4}$ (~-50 to 0‰)
444 (Etiope and Sherwood Lollar, 2013). In order to test this hypothesis, we performed
445 reconnaissance $\delta^{13}\text{C}$ analysis of CH₄ in the fluid inclusions (Boutier et al., in preparation). The
446 analyses yielded $\delta^{13}\text{C}$ in the range of -14‰ ($\pm 2\%$) for inclusions in olivine, and -13‰ ($\pm 1\%$)
447 for carbonate-hosted methane-rich inclusions from the rodingite. Even though these results

448 must be considered as preliminary test data, they seem to exclude the possibility of a pure
449 thermogenic source.

450 Following the hypothesis of an abiotic origin, two possible mechanisms for the formation of
451 CH₄ can be considered with either external or internal sources, respectively. External sources
452 correspond to the infiltration of CH₄ and other reduced species formed abiotically in other
453 geological reservoirs. A deep mantle origin for the reduced fluids detected in the BMC fluid
454 inclusions appears unlikely if the syn-inclusion, water-rich nature of the serpentinizing fluid is
455 considered. The metasedimentary formations adjacent to the BMC contain carbonate,
456 graphitic carbon, sulphides, and phyllosilicates as potential sources of the C-N-S-H fluid
457 identified in the fluid inclusions. Previous studies along the Appalachian belt have reported
458 evidence of carbon mobilization from these metasedimentary formations or their along-strike
459 equivalents. For example, Zhang *et al.*, 2018 documented metamorphic loss of isotopically
460 light carbon from the Wepawaug schists, Connecticut. The possibility for these carbonate-
461 bearing formations to generate strongly reduced fluids abiotically is not obvious —for
462 reference, water-maximum conditions in graphite-saturated fluids contain roughly equal
463 proportions of CH₄ and CO₂ (Connolly, 1995; Holloway, 1984) —. Nevertheless, evidence for
464 the circulation of CH₄-rich fluids in equivalent formations in East Central Vermont and New
465 Hampshire has been reported (Evans et al., 2002; Rumble III and Hoering, 1986).
466 Nevertheless, most petrological studies focusing on New England metasediment-derived
467 fluids point to more oxidized, CO₂-dominated aqueous fluids (Ferry, 2007; Penniston-Dorland
468 and Ferry, 2006). Alternative external sources of reduced fluids would require unidentified
469 processes, including mixing of different carbon reservoirs, or water-rock interactions
470 equivalent to those that took place in the BMC.

471 A reducing potential internal to the BMC, and in particular the hydroxylation of fresh mantle
472 peridotites during the Taconic subduction, appears the most likely interpretation for the
473 genesis of the identified reduced fluid species. The presence of Fe-Ni alloys in the BMC
474 partially serpentinized peridotites indicates that the rock recorded reducing conditions, as
475 already observed in several oceanic and ophiolitic, and some subduction zone serpentinites
476 (Evans et al., 2017; Frost, 1985; Klein and Bach, 2009; Vitale Brovarone et al., 2020). High-
477 pressure syn-serpentinization reducing conditions are also suggested by the thermodynamic
478 modelling results presented in this study, which indicates fO_2 values as low as $-3.2 \Delta QFM$ at
479 $400^\circ C$ and 1 GPa to $-1.2 \Delta QFM$ at $450^\circ C$ and 1 GPa, and the formation of significant amounts
480 of H_2 in the fluid (Fig. 10C). The interaction of this H_2 with dissolved carbon, nitrogen, and
481 sulphur species present in the serpentinizing fluid, could have favoured the formation of CH_4
482 H_2S , and NH_3 from more oxidized species. For NH_3 , the modelling also indicates that this
483 species is already the dominant N species in the infiltrating fluid buffered at QFM. This feature
484 suggests that the N_2 detected in the fluid is most likely formed through post-entrapment
485 respeciation of NH_3 , unless the serpentinizing fluid was more oxidized than QFM. The absence
486 of detectable H_2O in the fluid inclusions is interpreted to result from host-inclusion interaction
487 and formation of step-daughter lizardite and brucite, or by the preferential entrapment of
488 immiscible reduced gases relative to aqueous fluids (Huang et al., 2017; Vitale Brovarone et
489 al., 2017). The absence of residual H_2 in the fluid inclusions may be explained by the much
490 faster diffusion H_2 relative to other fluid species through the host olivine, or by selective
491 leakage.

492 The most plausible sources of serpentinizing fluid were the metasedimentary formations
493 surrounding the BMC complex. These rock types host substantial amounts of subducted
494 carbon, sulphur, and nitrogen (Bebout and Fogel, 1992; Evans et al., 2014; Kelemen and

495 Manning, 2015; Plank and Manning, 2019). An ultramafic source internal to the BMC would
496 not be consistent with the general retention of N during prograde metamorphism and
497 dehydration of serpentinites (Halama et al., 2014). The hypothesis of a metasediment-derived
498 serpentinizing fluid was also suggested by previous oxygen and hydrogen isotopic data on
499 antigorite from the BMC complex (Wenner and Taylor, 1974, 1971). Moreover, the authors
500 proposed antigorite-magnetite equilibration T in the range of 220-460 °C, which is consistent
501 with the prograde P-T of the BMC (Fig. 12). This supports the interpretation of a subduction-
502 related serpentinization related to the infiltration of metasedimentary-derived fluids in a
503 rather dry ultramafic body.

504 **5. Conclusions**

505 Mineralogical, microstructural, and fluid inclusion study of variably serpentinized dunite from
506 the Belvidere Mountain Complex, Appalachian belt, Northern Vermont, provides insight into
507 the process of high-pressure serpentinization in subduction zone and the related fluid-rock
508 redox patterns. Although the BMC underwent a complex tectonic evolution potentially
509 characterized by multiple stages of hydration from the (sub)seafloor to subduction and
510 exhumation, the collected data point to a major event of high-pressure serpentinization that
511 took place in the Appalachian subduction zone. Metamorphic olivine in the BMC
512 serpentinized peridotite is interpreted as the product of rock hydration rather than
513 dehydration, linked to the high-pressure serpentinization event. Our data support the
514 hypothesis that the methane observed in olivine-hosted secondary fluid inclusions is
515 genetically linked to serpentinization in the antigorite stability field consistent with the high-
516 pressure portion of the prograde or retrograde P-T path of the BMC. This favors an abiotic
517 origin for this methane, even though a biotic origin for the C source cannot be excluded. The

518 C- and N-rich composition of fluid inclusions suggests a metasediment-derived origin for the
519 serpentinizing fluid. The high-pressure serpentinization and related abiotic methanogenesis
520 place the Belvidere mountain complex as a suitable proxy for the study of mantle wedge
521 serpentinization. This study confirms the importance of fluid mobility in deep seated
522 ultramafic body in subduction zones for the mobility of C, H, and N, and their implications on
523 large-scale geochemical cycling.

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764

1 **Fig. 1:** A: Simplified geological map of Vermont, modified from Hibbard et al., (2006). B:
2 Simplified bedrock geologic map of the Belvidere Mountain Complex and the surrounding
3 formations. Modified after Hibbard et al., (2006). Units description from Hibbard et al., (2006)
4 and Gale, (2007).

5 **Fig. 2:** A-B: Photographs of the Belvidere serpentized peridotite in outcrop. C: Weakly
6 serpentized dunite (samples V18-2a and V18-2b). D: Serpentized dunite, with boudinaged
7 meta-gabbro (samples V18-3a and V18-3b).

8 **Fig. 3:** A: Photomicrograph of a partially serpentized dunite. B: Partial replacement of
9 primary chromian spinel by magnetite and chlorite. C: Partial replacement of olivine by
10 antigorite. D: Inferred primary orthopyroxene being pseudomorphically replaced by
11 metamorphic olivine. See also Figure 5B. E: Photomicrograph of a pyroxenite layer included in
12 the Belvidere peridotite. The photomicrograph shows radial diopside aggregate
13 pseudomorphic on primary clinopyroxene. Partially serpentized olivine is also visible. F:
14 Magnetite-rich diopside pseudomorphosis on primary clinopyroxene. Note the growth of
15 metamorphic olivine at the rim of the clinopyroxene site. See Figure 5A for SEM-based
16 backscattered electron image. G: Metamorphic diopside replacing primary clinopyroxene. In
17 this case, note the growth of metamorphic olivine along fractures cutting the
18 pseudomorphosis. See Figure 5F for SEM-based backscattered electron image. H: Relict of
19 metamorphic olivine pseudomorphic on primary clinopyroxene. The metamorphic olivine is
20 then partially replaced by serpentine + brucite + magnetite. A-D: Sample V18-2b, E-G: Sample
21 V18-3a; H: sample V18-B3. P-Ol: primary olivine, M-Ol: metamorphic olivine, Atg: antigorite,
22 Chl: chlorite, Mgt: magnetite, Di: diopside, Chr: chromite, Br: brucite, Ctl: chrysotile.

23 **Fig. 4:** Raman spectra of solid phases. A: Primary olivine. B Antigorite overgrowing primary
24 olivine. C: Antigorite core (Atg1) and rim (Atg2) (see Fig. 5A). D: Brucite.

25 **Fig. 5:** SEM-BSE images of samples V18-2b, V18-3a and V18-B3. A: Microstructural patterns of
26 serpentinization. Two generations of serpentine can be observed based on the BSE contrast, a
27 bright core (Atg1) and a dark rim (Atg2). Note also the striped zonation of primary olivine and
28 the formation of metamorphic olivine. B: Replacement of an inferred primary orthopyroxene
29 crystal by metamorphic olivine. The close-up shows the preservation of clinopyroxene relicts
30 interpreted as exsolutions inside the former orthopyroxene. C: Formation of at the expense of
31 metamorphic. D: Composite aggregate of Fe-Ni and Fe-Cu-Ni alloys and Ni sulphide. E: Growth
32 of metamorphic olivine around a diopside-rich primary clinopyroxene pseudomorphosis. A
33 primary chromina spinel partially converted into magnetite can also be observed. F:
34 Metamorphic diopside replacing primary clinopyroxene. Note the presence of antigorite +
35 metamorphic olivine ± magnetite along the fractures. In this case, metamorphic olivine
36 occupies the centre of the fractures and is not in contact with diopside, whereas the
37 metamorphic olivine rimming the primary clinopyroxene site is in contact with it. G:
38 Metamorphic olivine growing at the expense of metamorphic diopside (former primary
39 clinopyroxene). The microstructure suggests the former presence of serpentine needles
40 replacing the clinopyroxene and successively replaced by metamorphic olivine. Both Atg1 and
41 Atg2 antigorite generations are present. H: Relict of metamorphic olivine formed at the
42 expense of a primary orthopyroxene site in V18-3b. The metamorphic olivine is partially
43 converted into brucite + serpentine. P-Ol: primary olivine M-Ol: metamorphic olivine Px:
44 pyroxene Di: diopside Mtg: magnetite Chr: chromite Atg: antigorite Br: brucite Ctl: chrysotile
45 NiFeS: nickel and iron sulphite. The presence of antigorite was confirmed by Raman
46 spectroscopy.

47 **Fig. 6:** Mg# versus MnO (wt%) diagram showing the compositional variation of primary and
48 metamorphic olivine. The Mn-richest cluster of metamorphic olivine belongs to sample V18-
49 B3 (fully serpentinized peridotite). Background data from Arai et al., (2012); Dandar et al.,
50 (2019); Debret et al., (2013); Iyer et al., (2008); Nozaka, (2018); Plümper et al., (2012b);
51 Scambelluri et al., (1995); Shen et al., (2015).

52 **Fig. 7:** Photomicrographs of methane-rich fluid inclusion trails in olivine from sample V18-2a.
53 Black arrows indicate fluid inclusion trails. A: Inclusion trail being cut by antigorite veinlets. B:
54 Inclusion trail limited by antigorite veins. C: Secondary trail of fluid inclusions propagating
55 from an antigorite veinlet. D: Photomicrograph showing a fluid-inclusion-rich olivine
56 aggregate. E: Close up of the methane-rich fluid inclusions. F: Close up of a graphite bearing
57 fluid inclusion (red arrow), as confirmed by Raman spectroscopy in (Fig. 8B).

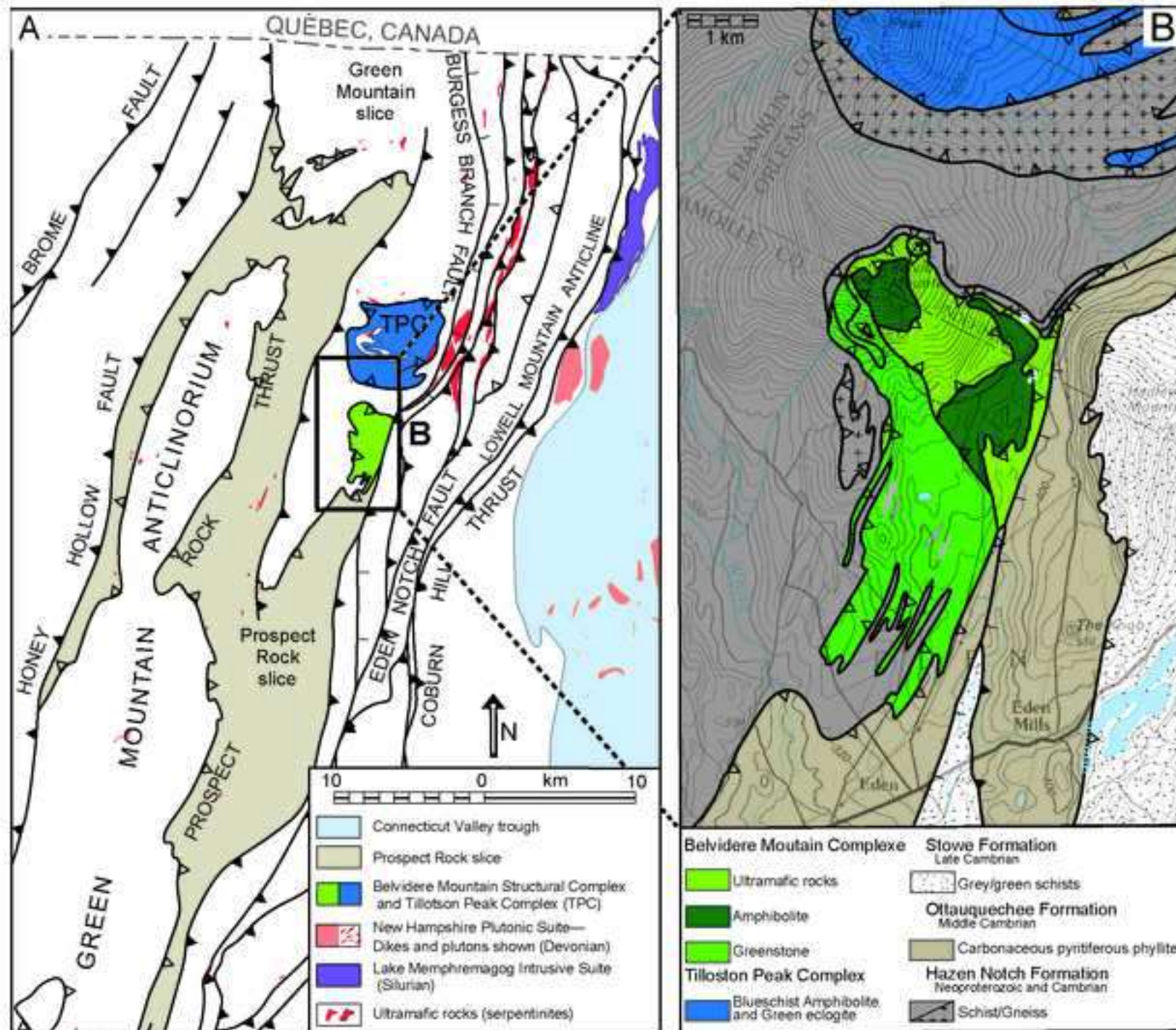
58 **Fig. 8:** Raman spectra of fluid inclusions and step-daughter solid phases. A: Inclusion showing
59 a marked CH₄ band and minor peaks of N₂, NH₃ and S-H bond. Lizardite and brucite O-H bands
60 are also observed. B: Graphite in fluid inclusions (see Fig. 6D). C: Methane-rich fluid inclusion
61 with a close up of O-H bonds of lizardite and brucite.

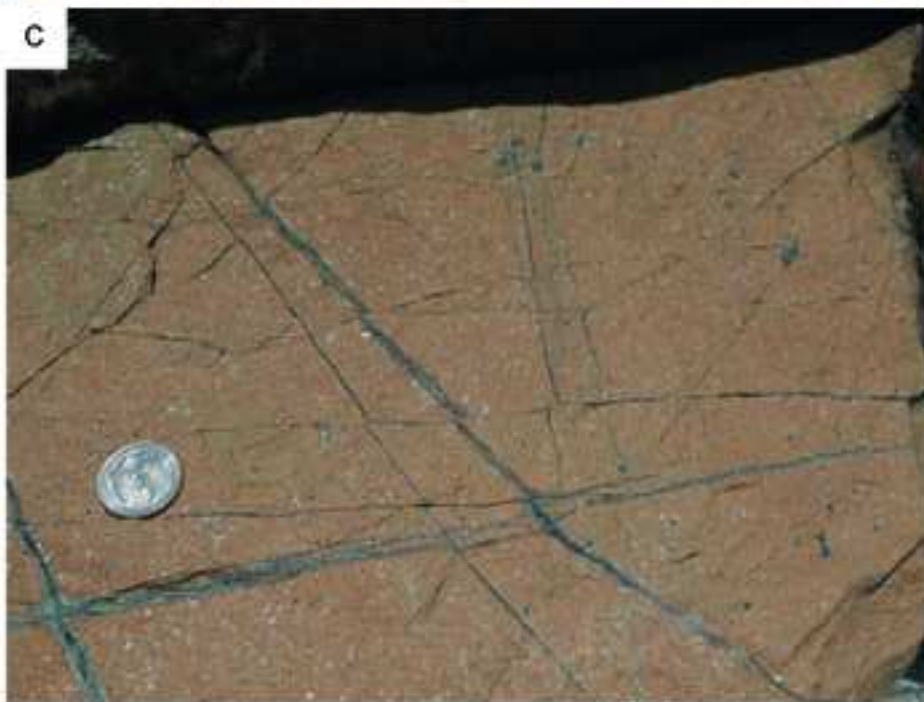
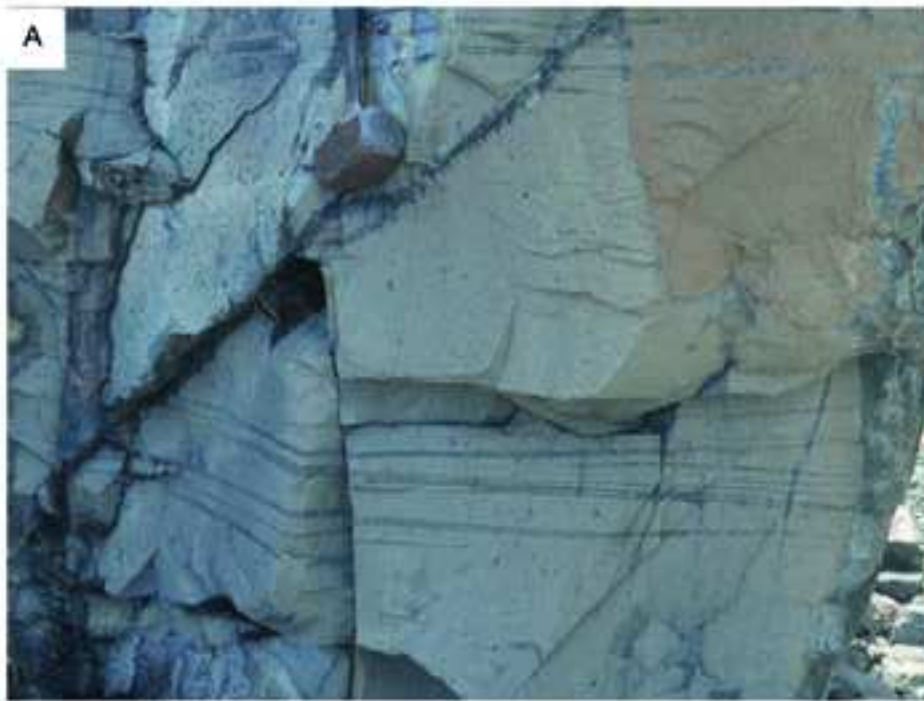
62 **Fig. 9:** Calculated molar fraction of CH₄, N₂, H₂S and NH₃ from Raman scattering cross-section
63 and the instrumental efficiency of each species. See Frezzotti et al., (2012) for methodology.

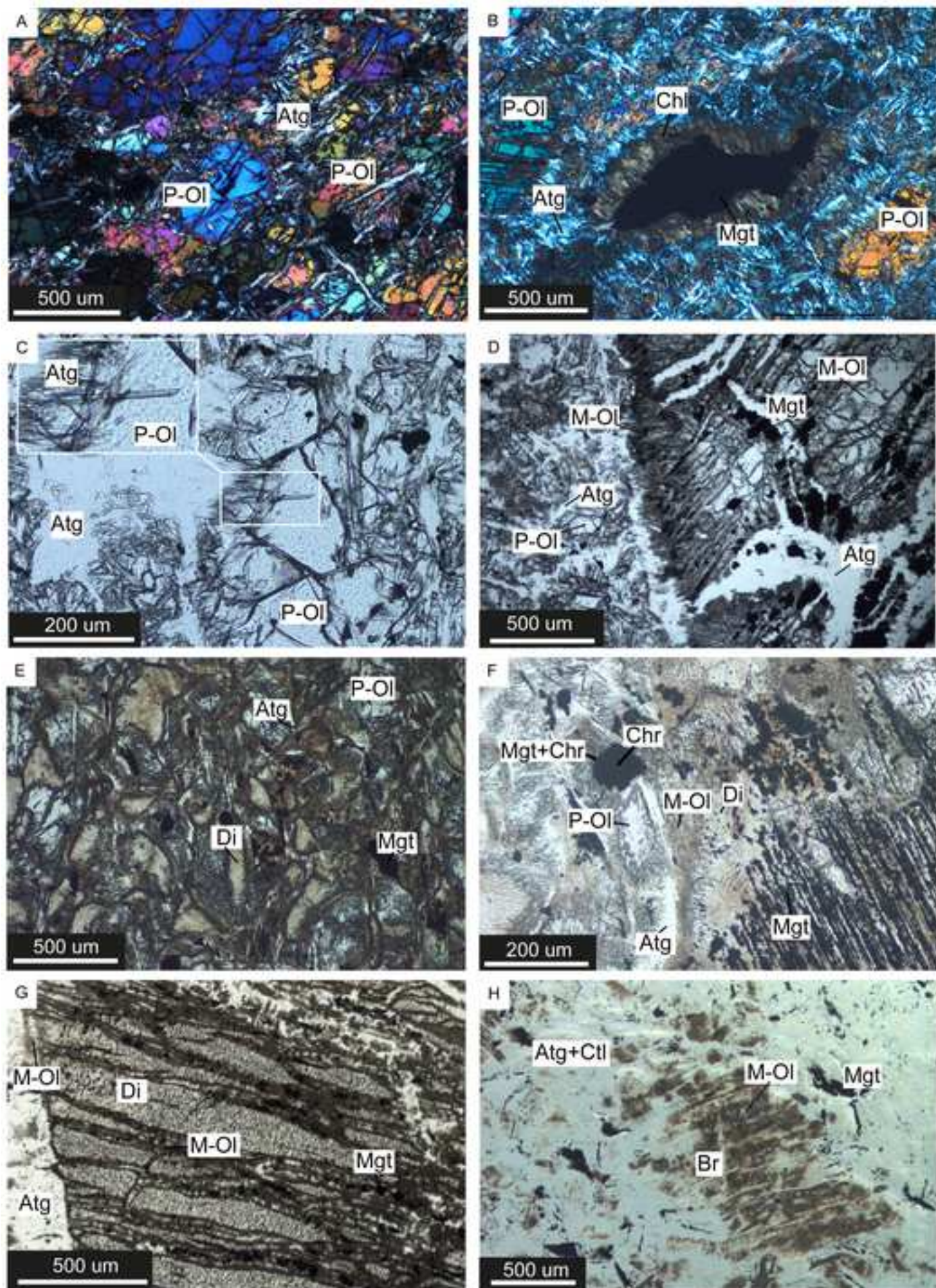
64 **Fig. 10:** Thermodynamic modelling of HP serpentinization of dunite and related mineralogical
65 and fluid evolution. A: Mineralogical evolution during serpentinization at 450 °C and 1 GPa.
66 Representative microstructures of the main reaction steps are proposed as observed in the
67 natural samples. B: Evolution of the nitrogen and sulphur fluid speciation as a function of the
68 reaction progress presented in A. C: Evolution of the fO₂ and H₂, CH₄, and CO₂ concentrations
69 in the fluid as a function of the reaction progress presented in A.

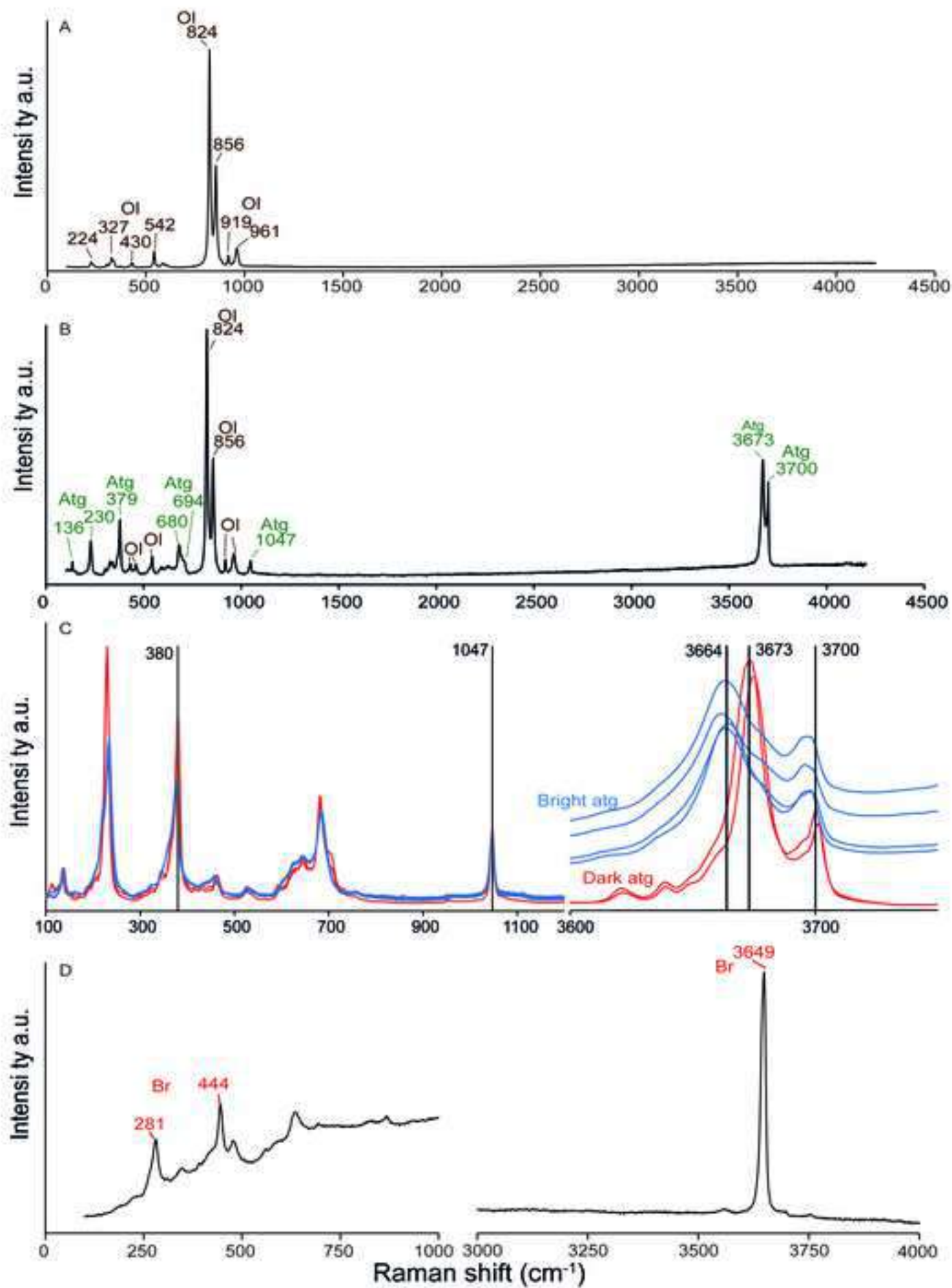
70 **Fig. 11:** Reconstruction of the mineralogical evolution of the BMC partially serpentized
71 peridotite. An early step of serpentization is proposed based on the needle-like growth of
72 metamorphic olivine on primary clinopyroxene, suggesting the presence of serpentine prior to
73 the formation of metamorphic olivine. The successive growth of antigorite at the expense of
74 both primary and metamorphic olivine constrains the main serpentization event to
75 metamorphic conditions. Finally, a late serpentization event is proposed based on the
76 growth of brucite + antigorite + chrysotile on relict metamorphic. Chr : chromite, P-Ol : primary
77 olivine, M-Ol : metamorphic olivine, Opx : primary orthopyroxene, Cpx : primary clinopyroxene,
78 Chl : chlorite, Mgt : magnetite, Atg : antigorite, Di : diopside, Brc : brucite, Ctl : chrysotile.

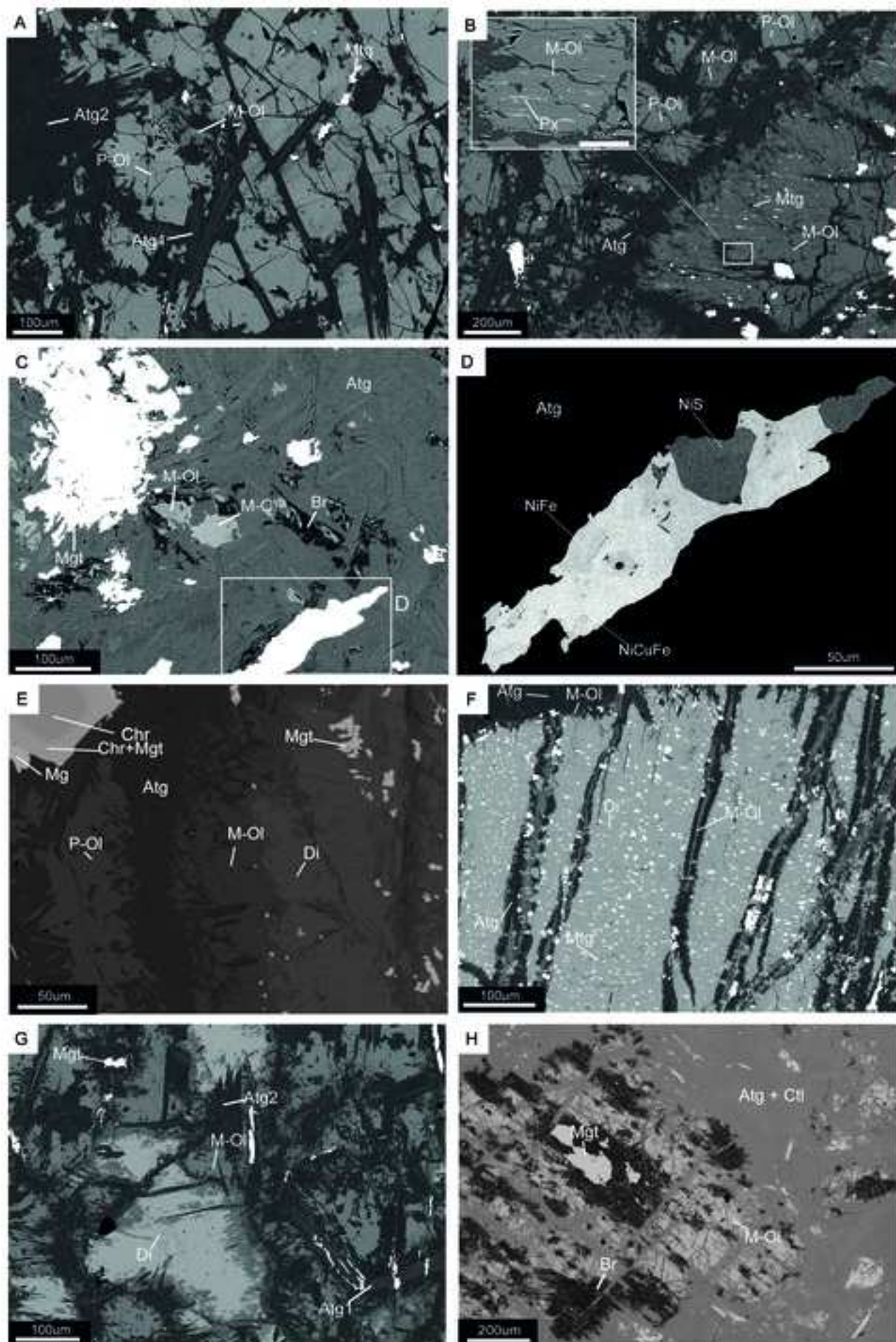
79 **Fig. 12:** Stability field of serpentine type minerals and olivine, modified from Guillot et al.,
80 (2015) (see references therein for details on the main reactions). The retrograde P-T path of
81 BMC from Honsberger, (2015) is also shown for reference. Atg : antigorite, Brc : brucite; Ctl :
82 chrysotile; Ilm : ilmenite; Ol : olivine; Tlc : talc; Ticl : titanian clinohumite.

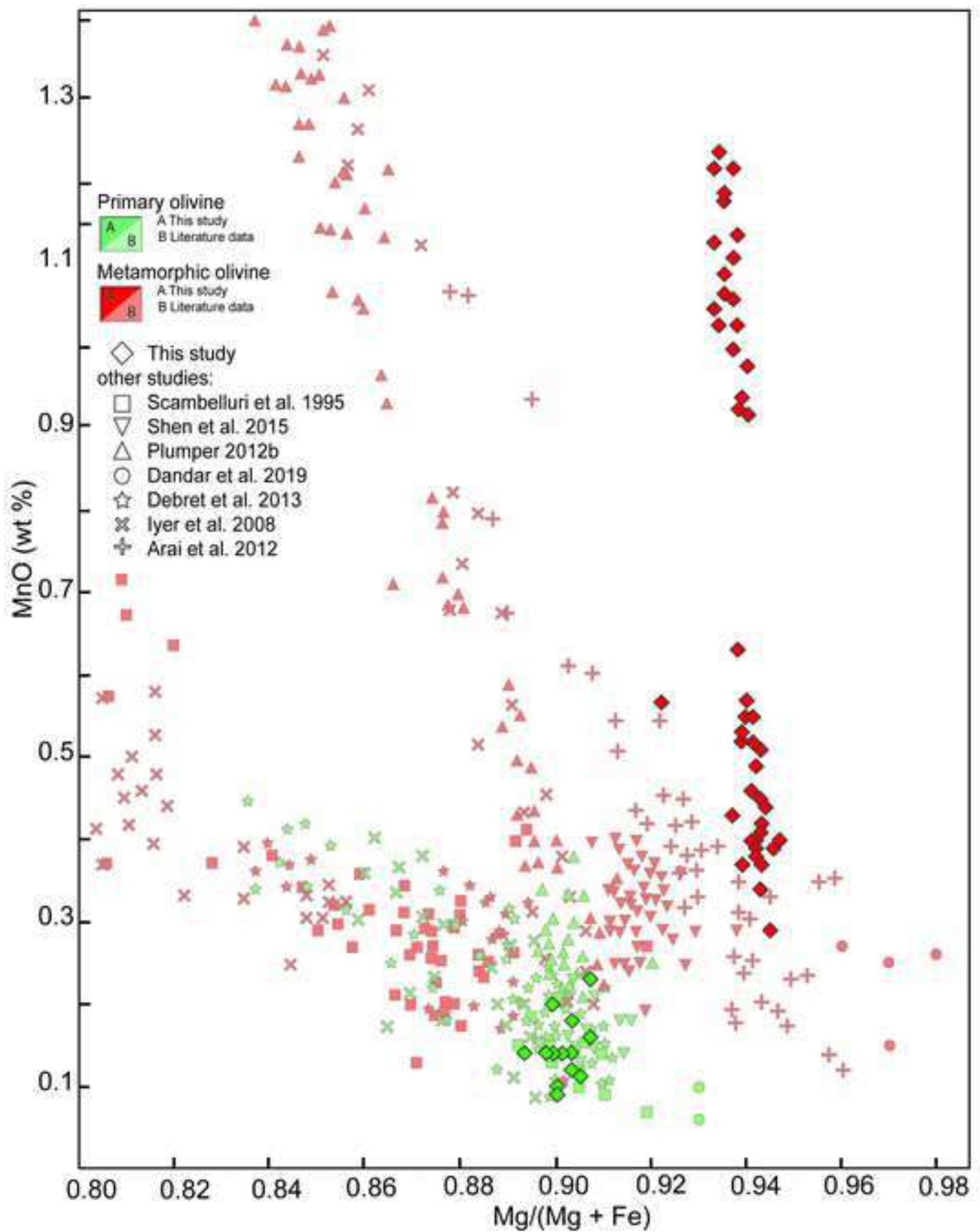


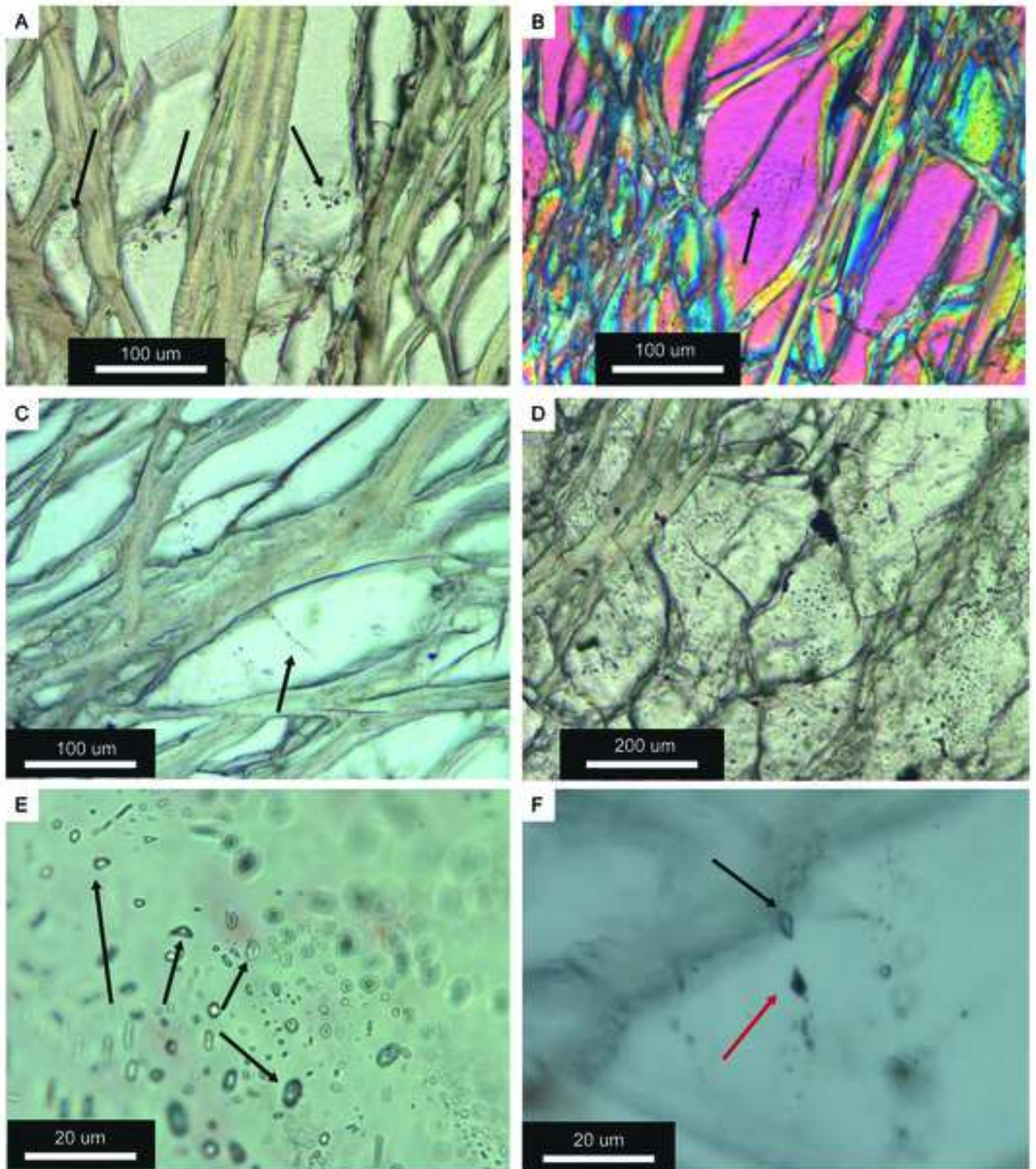


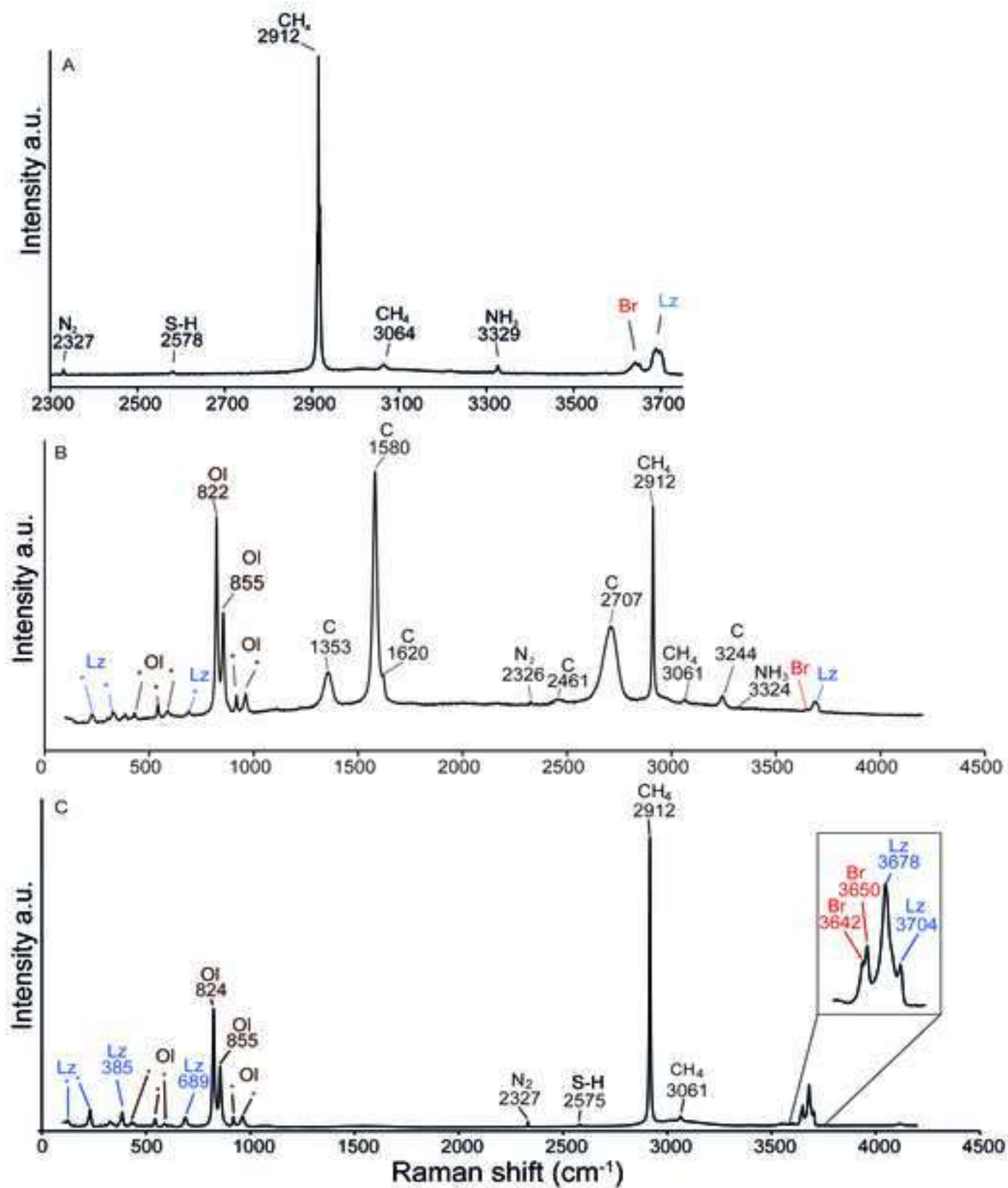


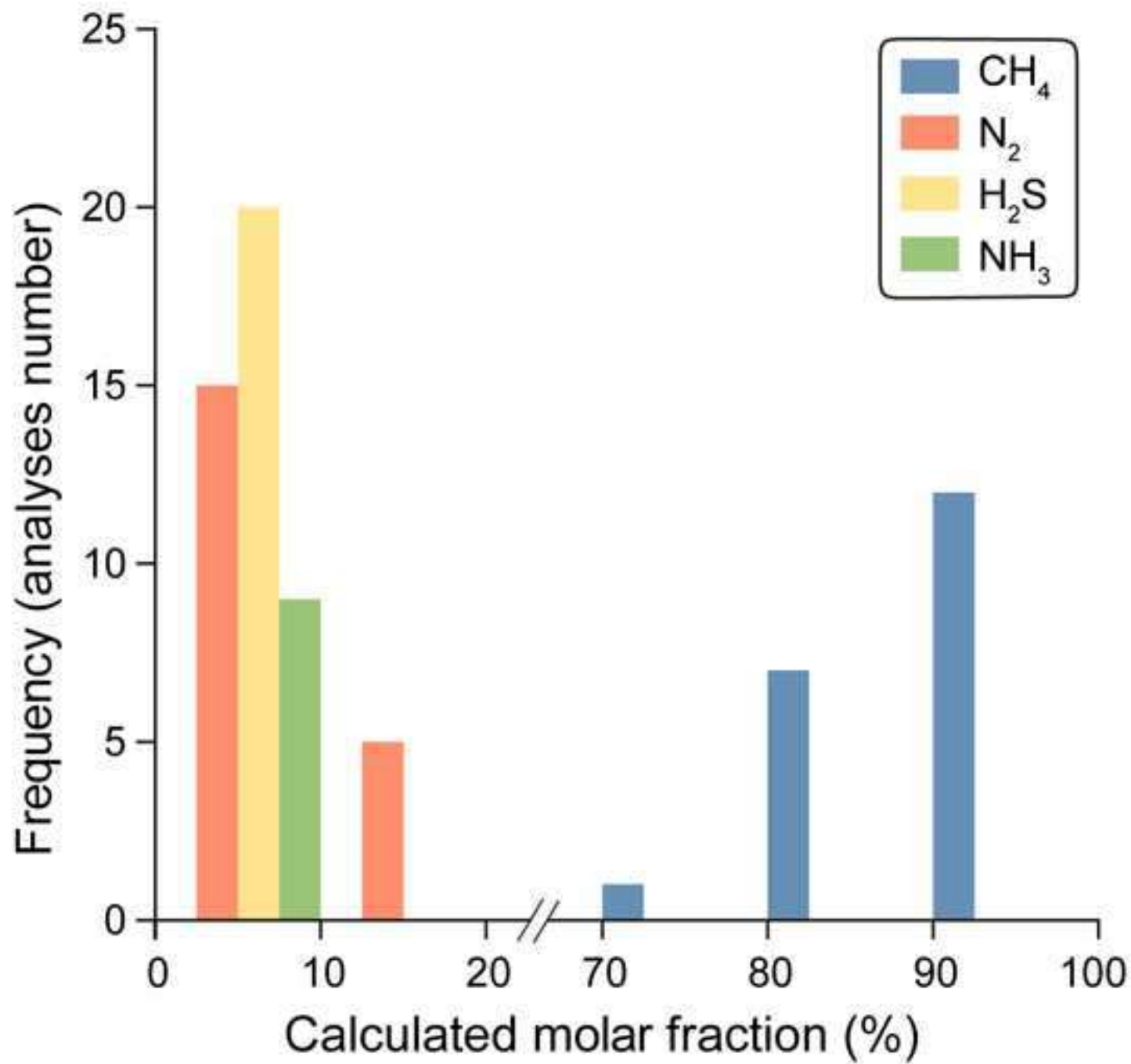


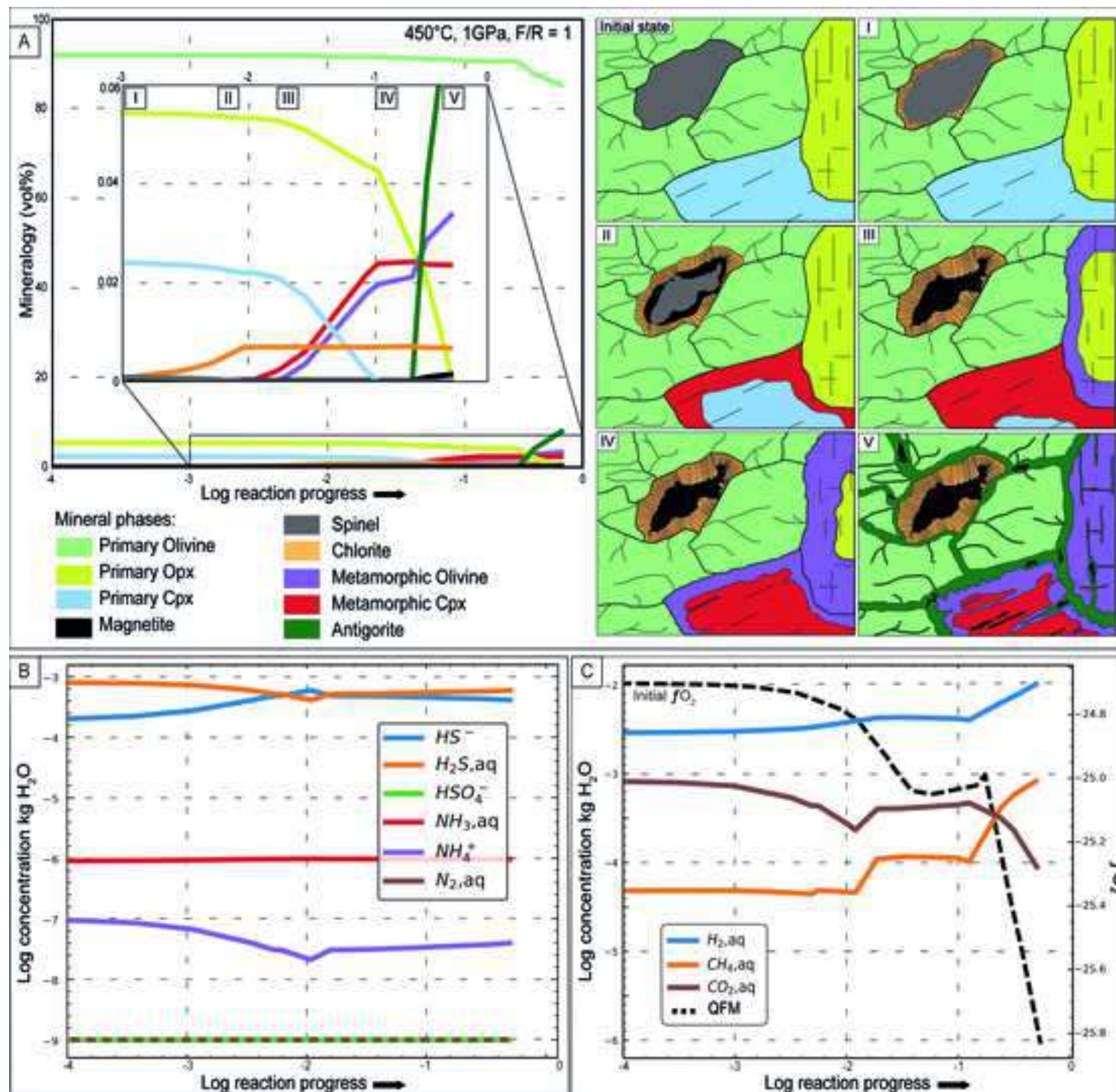


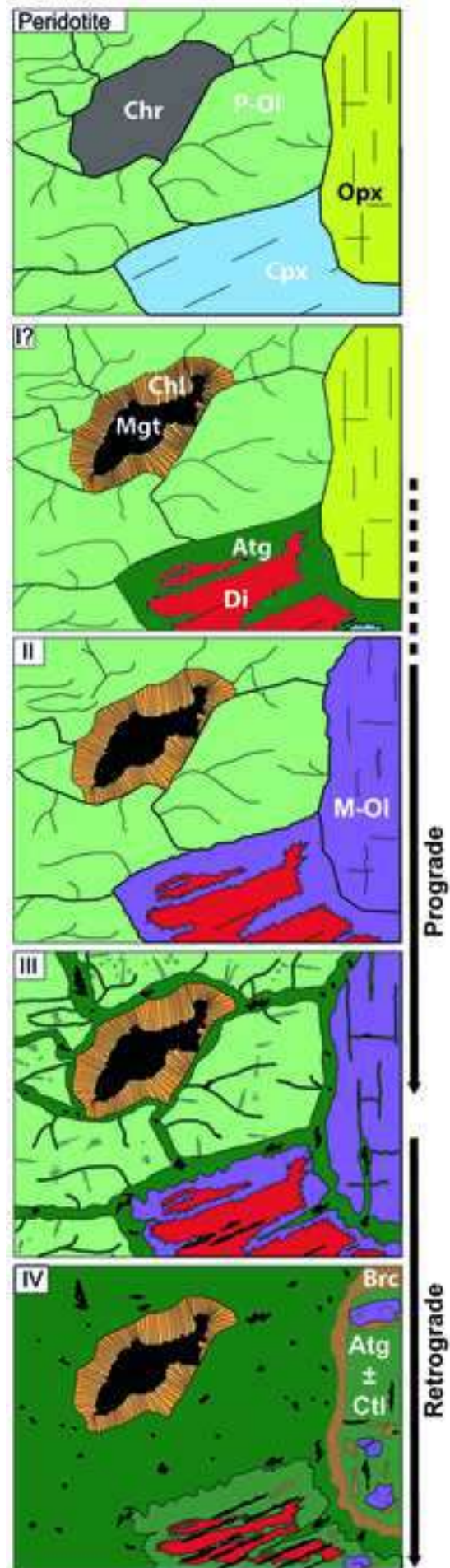


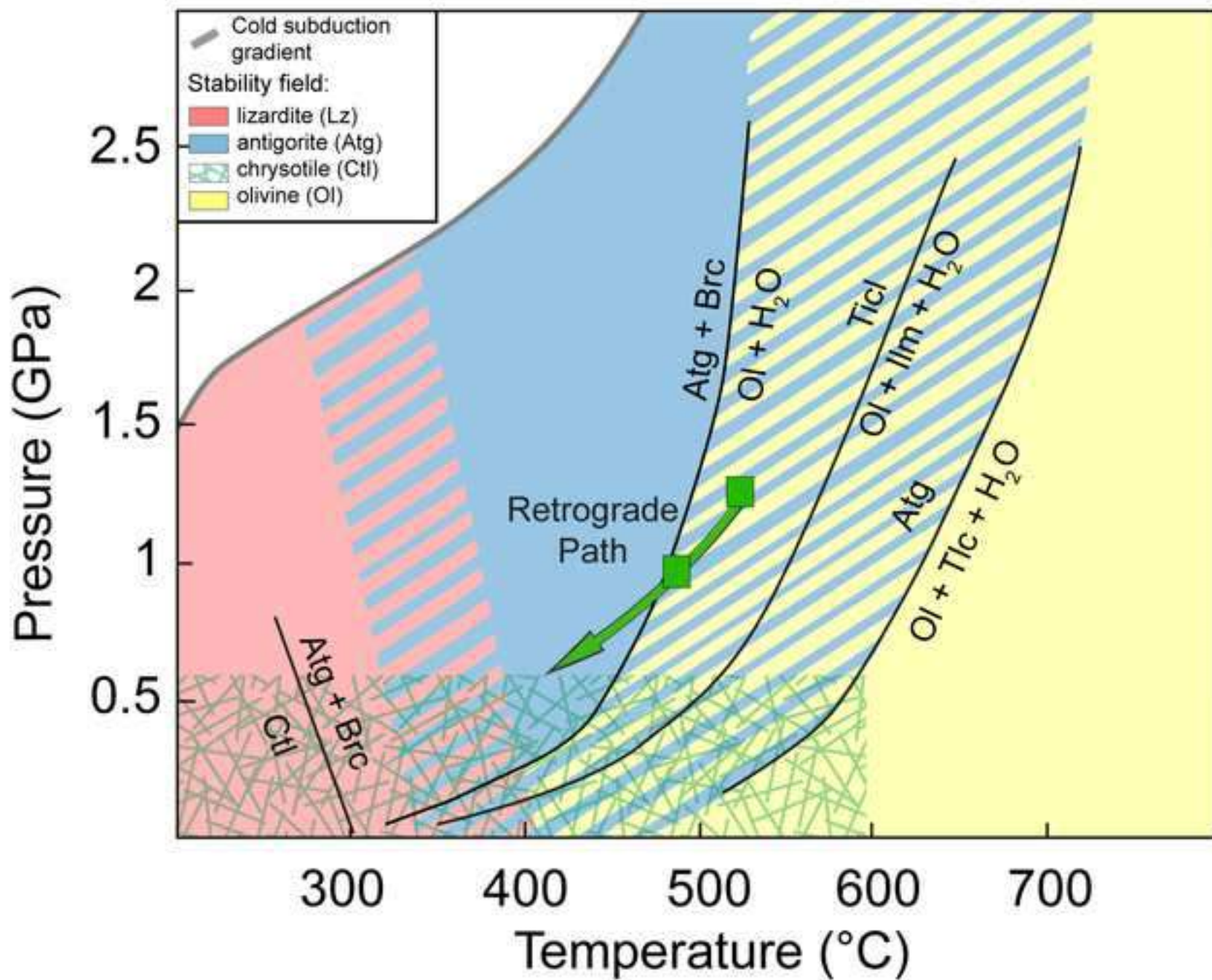












1 **Table 1**

MicroProbe analysis. Absence of SD value when n>1 indicates SD<0.01. *EDS analyses normalized at 100.

Minerals	Primary Olivine		Metamorphic Olivine			Pyroxene	Serpentine (Bright core)		Serpentine (Matrix)		Chlorite	
	V18-2b	V18-3a	V18-2b	V18-3a	V18-B3*	V18-3a	V18-2b	V18-3a	V18-2b	V18-3a	V18-2b	V18-3a
Sample (n)	14	11	28	15	25	20	6	2	9	4	11	1
SiO ₂	40.91 (0.25)	40.83 (0.46)	41.31 (0.49)	41.95 (0.63)	42.00 (0.12)	54.96 (0.98)	42.87 (1.21)	43.41 (0.1)	44.15 (0.17)	43.1 (1.35)	32.61 (1.34)	32.03
TiO ₂	0.01 (0.01)	0.02 (0.02)	0.02 (0.01)	0.02 (0.02)		0.03 (0.02)	0.01 (0.01)	0.02 (0.01)	0.02 (0.01)	0 (0)	0.02 (0.02)	0.00
Al ₂ O ₃	0.01 (0.01)	0.01 (0.01)	0.01 (0.02)	0.04 (0.09)		0.15 (0.11)	1.77 (1.22)	1.45 (0.78)	0.52 (0.21)	0.47 (0.26)	14.34 (2.39)	11.78
Cr ₂ O ₃	0.01 (0.02)	0.01 (0.02)	0.06 (0.18)	0.04 (0.03)		0.17 (0.11)	0.61 (0.85)	0.25 (0.19)	0.13 (0.09)	0.1 (0.1)	1.07 (0.56)	4.07
FeO (tot)	8.97 (0.34)	7.93 (0.22)	4.83 (0.30)	5.05 (0.57)	4.71 (0.19)	0.76 (0.16)	2.60 (0.26)	1.58 (0.27)	1.35 (0.22)	1.2 (0.29)	3.31 (0.46)	2.79
MnO	0.15 (0.04)	0.12 (0.06)	0.45 (0.08)	0.37 (0.06)	1.02 (0.18)	0.05 (0.03)	0.04 (0.02)	0.05 (0.05)	0.04 (0.02)	0.02 (0.02)	0.06 (0.10)	0.07
NiO	0.38 (0.03)	0.49 (0.06)	0.43 (0.07)	0.39 (0.06)	0.46 (0.08)	0.03 (0.03)	0.16 (0.02)	0.18 (0.06)	0.15 (0.04)	0.19 (0.02)	0.22 (0.05)	0.24
ZnO	0.04 (0.04)	0.03 (0.04)	0.03 (0.04)	0.03 (0.04)		0.01 (0.02)					0.02 (0.04)	0.00
MgO	50.23 (0.22)	50.91 (0.22)	52.8 (0.64)	51.8 (1.57)	51.80 (0.28)	18.75 (1.15)	37.97 (1.30)	39.81 (0.59)	39.6 (0.77)	40.07 (0.61)	34.34 (1.11)	35.08
CaO	0.02 (0.01)	0.01 (0.02)	0.02 (0.02)	0.25 (0.37)		24.87 (1.43)	0.02 (0.02)	0.05 (0.04)	0.02 (0.02)	0.02 (0.01)	0.02 (0.01)	0.05
Na ₂ O	0.09 (0.22)	0.01 (0.01)	0.02 (0.02)	0.01 (0.01)		0.04 (0.02)					0.02 (0.03)	0.03
K ₂ O	0.01 (0.02)	0.01 (0.01)	0.01 (0.01)	0.01 (0.01)		0.01 (0.01)					0.02 (0.02)	0.02
Total,	100.82 (0.36)	100.37 (0.70)	100.06 (1.10)	99.96 (1.63)	100	99.83 (1.12)	86.2 (0.63)	86.93 (0.4)	86.07 (0.82)	85.25 (1.66)	86.14 (0.65)	86.17
Cations												
Si	0.990	0.988	0.991	1.012	1.013	1.989	1.984	2.010	2.044	1.995	6.232	6.165
Ti	0.000	0.000	0.000	0.000		0.001	0.000	0.001	0.001	0.000	0.004	0.000
Al	0.000	0.000	0.000	0.001		0.006	0.091	0.079	0.028	0.026	3.229	2.672
Cr	0.000	0.000	0.001	0.001		0.005	0.026	0.009	0.005	0.004	0.162	0.619
Fe ²⁺ (tot)	0.181	0.161	0.097	0.102	0.095	0.023	0.104	0.061	0.052	0.046	0.529	0.404
Mn	0.003	0.003	0.009	0.008	0.021	0.002	0.001	0.002	0.002	0.001	0.010	0.012
Ni	0.007	0.010	0.008	0.008	0.009	0.001	0.006	0.007	0.006	0.007	0.034	0.037
Zn	0.001	0.001	0.000	0.001		0.000					0.003	0.000
Mg	1.812	1.837	1.891	1.861	1.862	1.012	2.486	2.747	2.733	2.765	9.782	10.063
Ca	0.000	0.000	0.001	0.006		0.964	0.001	0.002	0.001	0.001	0.003	0.011
Na	0.004	0.000	0.001	0.000		0.003					0.007	0.012
K	0.000	0.000	0.000	0.000		0.000					0.005	0.005
Mg#	0.91	0.92	0.95	0.95	0.95	0.98	0.94	0.96	0.97	0.97	0.95	0.96

Mg# = Mg/(Mg+ΣFe)

3 **Table 2**

MicroProbe analysis. Absence of SD value when n>1 indicates SD<0.01

Minerals	Spinel (nucleus)	Spinel (intermediate)	Magnetite (rim)	Magnetite (mesh and veins)	Brucite	
Sample (n)	V18-3a	V18-3a	V18-3a	V18-3a	V18-2b	V18-2b
	7	7	7	1	4	10
SiO ₂	0.02 (0.02)	0.03 (0.01)	0.06 (0.07)	0.07	0.05 (0.03)	0.78 (2.20)
TiO ₂	0.09 (0.02)	0.23 (0.05)	0.03 (0.02)	0.04	0.01 (0.01)	0.01
Al ₂ O ₃	12.63 (0.94)	3.87 (1.34)	0.00	0.03	0.01 (0.01)	0.01 (0.01)
Cr ₂ O ₃	47.51 (0.57)	41.6 (1.29)	1.87 (1.16)	0.01	0.02 (0.02)	0.04 (0.05)
FeO	34.12 (1.1)	48.32 (2.02)	91.45 (0.75)	92.44	92.57 (0.28)	2.91 (0.18)
MnO	0.34 (0.04)	0.95 (0.36)	0.16 (0.03)	0.12	0.19 (0.05)	0.22 (0.04)
NiO	0.09 (0.02)	0.27 (0.05)	1.05 (0.1)	1.11	0.82 (0.09)	0.38 (0.09)
ZnO	0.49 (0.08)	0.3 (0.1)	0.04 (0.06)	0.10	0.04 (0.08)	0.01 (0.02)
MgO	5.99 (0.32)	3.37 (0.22)	0.71 (0.09)	0.62	0.82 (0.19)	75.88 (3.93)
CaO	0.01 (0.01)	0.00	0 (0.01)	0.00	0.00	0.02 (0.03)
Na ₂ O	0.03 (0.04)	0.02 (0.02)	0.03 (0.03)	0.03	0.02 (0.04)	0.02 (0.02)
K ₂ O	0.00	0.01 (0.01)	0 (0.01)	0.01	0.00	0.01 (0.01)
Total	101.4 (0.56)	99.03 (0.6)	95.46 (0.5)	94.63	94.56 (0.45)	80.33 (2.70)
Cr#	0.79(0.01)	0.92(0.2)	1.00			

4 Cr# = Cr/(Al+Cr)

5 **Table 3**

MicroProbe analysis of alloys

Sample Alloy	V18-2b						V18-3a		
	NiFeCu	NiS	NiS	NiS	NiFeCu	NiFeCu	NiS	NiS	NiS
S		24.06	24.25	24.12			24.59	24.52	24.50
Fe	12.21	0.9	0.51	1.35	13.35	20.07	0.84		0.95
Pb				0.13			0.18		
Ni	86.23	72.68	72.71	73.55	86.05	80.03	72.97	72.85	73.30
Cu	1.94				1.75	1.49			
Total:	100.45	97.9	97.59	99.18	101.28	101.71	98.70	97.72	99.04

6 **Table 4**

Calculated partition coefficient between antigorite and olivine for Mg and Mn

Calculated K _D	K _D Mg		K _D Mn	
	V18-2	V18-3b	V18-2	V18-3b
Atg1 / P-OI	0.25	0.14	0.42	0.54
Atg2 / P-OI	0.11	0.11	0.42	0.27
Atg1 / M-OI	0.49	0.23	0.14	0.18
Atg2 / M-OI	0.21	0.17	0.14	0.09

K_D Atg/OI from bibliography

0.45-0.35

0.18

K_D Atg/OI Mg from Evans et al.(2008) and Mn from Trommsdorff and Evans, 1974.