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1 Pre and post failure dynamics of landslides in the Northern Apennines
2 revealed by space-borne synthetic aperture radar interferometry (InSAR)

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6 **Abstract**

7 Landslides are common landscape features in the Northern Apennine mountain chain and cause
8 frequent damages to human structures and infrastructure. Most landslides in the area can be
9 classified as earthflows, where the clay-shales form the substrate, whereas complex landslides with
10 flow and sliding components are common on the slopes where fine-grained turbidites form the
11 substrate. Most of these landslides move periodically with contained velocities and, only after
12 particular rainfall events, some of them accelerate abruptly. Space-borne synthetic aperture radar
13 interferometry (InSAR) provides a particularly convenient way for studying the periods before
14 and after failures. In this paper, we present InSAR-results derived from the Sentinel 1 satellite
15 constellation for two landslide cases in the Northern Apennines. The first case is a complex
16 landslide that is hosted on a pelitic flysch formation, whereas the second case is an earthflow
17 located in chaotic clay shales. Both cases failed catastrophically and threatened or damaged
18 important infrastructures. In the case of the complex landslide, we report spatially variations
19 of the deformation field between repeated periods of acceleration. The data illustrate that the
20 deformation initiated in the upper part of the slope and expanded over the whole landslide body
21 afterward. In the case of the earthflow, we describe spatial and temporal kinematics during the
22 period before a catastrophic failure in March 2018. We discuss the temporal deformation signal
23 together with rainfall and snowmelt data from a nearby meteorological station. Deformation and
24 precipitation data highlight that high total precipitation can be considered the trigger of the failure.

25 *Keywords:* InSAR, landslides, earthflows, failure, rainfall, snowmelt

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26 1. Introduction

27 Landslides are common morphological features throughout the whole Northern Apennines
28 chain. Most slope deformations occur on old landslide materials that failed in the past (Bertolini
29 et al., 2004). In many cases, the reactivation of old deposits causes the regression of the main scarp
30 and the physical degradation of the material which may move downwards as an earthflow. In other
31 cases, the reactivation is more complex and different types of landslides can occur (Bertolini and
32 Pellegrini, 2001). Because the typical velocity of most of these landslides can span from centimeters
33 per year to meters per hour, depending on the stage of life (Cruden and Varnes, 1996), it becomes
34 important to properly monitor the displacements of the involved masses to assess the possibility
35 of sudden accelerations.

36 A powerful technique for monitoring the displacements of large areas is the synthetic aperture
37 radar interferometry (InSAR) that provides the possibility to measure the deformations of the
38 landslide deposits during the slow-motion stage (i.e., before the rapid acceleration). InSAR was
39 applied in a landslide-prone area in the mid-1990s (Fruneau et al., 1996), but only in the 2000s
40 it became a well-known technique for landslide monitoring. The development of multi-temporal
41 methods (e.g. Ferretti et al., 2001; Berardino et al., 2002; Hooper et al., 2004; Hooper, 2008) helped
42 in many cases to obtain useful InSAR derived information on the displacement of landslides. Those
43 techniques have been developed to overcome some of the limitations that conventional two-pass
44 interferometry had shown until that time (Colesanti and Wasowski, 2006; Wasowski and Bovenga,
45 2014). Since then, different InSAR techniques have been used to retrieve spatial and temporal
46 deformations of landslide-prone slopes in soft rocks (Colesanti et al., 2003; Hilley et al., 2004;
47 Wasowski and Bovenga, 2014; Handwerger et al., 2015; Bayer et al., 2017, 2018).

48 The two most common multi-temporal techniques are the Persistent Scatterer interferometry
49 (PS-InSAR, Ferretti et al., 2001) and Small Baseline techniques (SBAS, Berardino et al., 2002,
50 Schmidt and Bürgmann, 2003): the former is based on the stable SAR response of specific targets
51 (i.e. stable scatters), computed by using single-master interferograms series; the latter is often
52 optimized to derive spatially distributed information of multi-master interferograms series. Other
53 techniques combine the advantages of both techniques (Hooper, 2008).

54 PS-InSAR and small baseline techniques are widely used for landslide studies (Bianchini et al.,
55 2013; Tofani et al., 2013; Wasowski and Bovenga, 2014; Raspini et al., 2019), but in mountainous

56 areas the quality of measurements is often affected by decorrelation from the environmental setting
57 and in particular the presence of snow during the winter months and vegetation in the rest of
58 the year. In such contexts, stable scatters detection is constrained to human structures which are
59 characterized by high coherence values. Thus, decorrelation issues are still challenging in scarce-
60 urbanized areas.

61 In the past, only L-band data delivered spatially quasi-continuous data in settings similar to
62 the northern Apennines. The few reported examples, however, resolve mainly on the seasonal
63 kinematics of slow-moving landslides in California (Roering et al., 2009; Handwerger et al., 2013).
64 The launch of the new Sentinel 1 satellite constellation, which is characterized by a high acquisition
65 frequency of up to six days, is suited to reduce decorrelation in the derived interferograms (Intrieri
66 et al., 2018; Carlà et al., 2018) and permits to obtain promising results with higher temporal
67 resolution (Handwerger et al., 2019).

68 In this paper, we investigate the response of two landslides using InSAR analysis. Because
69 the landslides are located in rural areas with scarce urbanization, we use standard InSAR and
70 explore its potential in capturing the changeable rates of displacement and spatial patterns of
71 deformation. The two landslides were selected because they experienced catastrophic failures
72 (here defined as stage A of the morphological classification reported in Picarelli et al. (2005);
73 failure in the following), during the time span of our investigation. These circumstances offer the
74 possibility to explore standard InSAR potential to detect pre- and post-failure deformations and
75 document its evolution through time. Though such documentation has been previously reported
76 for instrumented landslides (e.g., Scoppettuolo et al., 2020), the possibility to use InSAR implies
77 the advantages intrinsic to remote sensing techniques that open to applications that include areal
78 surveillance and early detection.

79 We show that the technique is capable of producing spatially quasi-continuous maps of defor-
80 mation, also in areas that are characterized by the absence of good quality reflectors. Our data
81 indicate that, in both cases, the failure was preceded by detectable deformation. InSAR results
82 document the pre-failure and post-failure stages of the movement in terms of its spatial pattern
83 and temporal evolution. In one of the two cases, we could derive actual displacement time-series
84 that were compared to the precipitation regime to identify the triggering condition.

85 **2. Geological and geographical background**

86 The northern Apennine mountain chain is a pile of thrust and nappe units, transported towards
87 the Padan-Adriatic-Ionian-Hyblean foreland starting from Late Oligocene times. In the Northern
88 Apennines, the most common lithologies are chaotic clay shales and flysch deposits (Royden et al.,
89 1987; Castellarin, 1992; Patacca et al., 1993; Marroni and Treves, 1998).

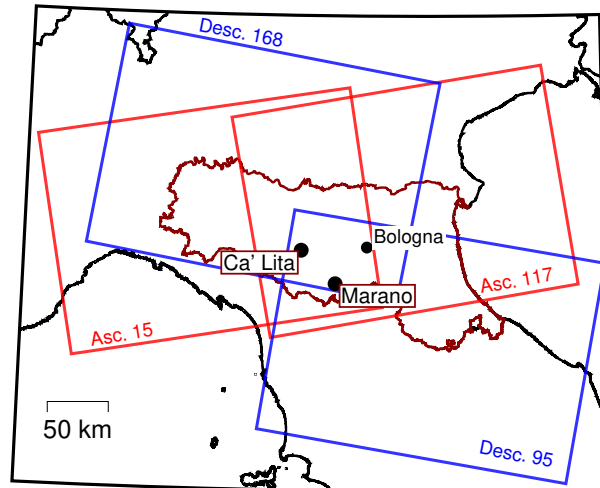


Figure 1: Location of the investigated landslides and setting of the satellite Tracks over the area. The region boundaries are highlighted by the dark red line and the landslides sites are labeled. The red boxes represent ascending Tracks and the blue boxes represent the descending Tracks.

90 The northern Apennines are affected by a high density of landslides and Bertolini and Pel-
91 legrini (2001) reported more than 32,000 landslides over the region of Emilia Romagna. In the
92 classification scheme of Cruden and Varnes (1996), most of them can be described as complex
93 landslides, associating roto-transitional slides with earthflows. Typical velocities are millimeters to
94 centimeters per year during the dormant phase (which may last years to hundreds of years) and
95 may increase up to meters per hour during the failure. The failure stages typically occur after
96 periods of large amount of rainfall. The average annual rainfall at elevations similar to those of
97 the two case studies is around 1200 mm, but the pluviometric regime is not uniform and 75% of
98 the total rainfall occurs in two rainfall seasons, one of which occurs during fall and one during
99 spring (Bertolini and Pellegrini, 2001; Berti and Simoni, 2012; Berti et al., 2012). The investigated
100 cases are located in the Northern Apennines of Italy and both of them are covered by four Sentinel
101 satellite orbits, two of which imaged the area in ascending geometry, whereas the other two swaths

102 cover in descending geometry (Fig 1). The landslides reached the failure in the period covered by
103 Sentinel 1 flights. Marano reactivated in March 2018 and Ca Lita in March 2016 and November
104 2017.

105 *2.1. The Ca Lita landslide*

106 The Ca Lita landslide (Fig. 1, 2 a) developed on a hillslope composed of flysch and clay-shales
107 belonging to the Ligurian Units (Papani et al., 2002). It is located between 230 and 640 m a.s.l.
108 in the Reggio Emilia province (Italy); the total length is 2.7 km, with a mean slope angle of 15
109 degrees and a total estimated volume of 40 Mm^3 . The landslide can be classified as a reactivated
110 complex landslide (Cruden and Varnes, 1996), in which a rotational rock slide in the head zones
111 (in the Monghidoro Flysch Formation) evolves into an earthflow in the lower main body (in the
112 Rio Cargnone Clayshales). It reactivated several times in the last century (Borgatti et al., 2006;
113 Corsini et al., 2006; Cervi et al., 2012).

114 One catastrophic failure occurred in early spring 2004 after an intense rainy and snowy period.
115 During this reactivation, it reached peak velocities of about 10 m per day at the toe and only
116 of few decimeters per day in the upper part (Borgatti et al., 2006; Corsini et al., 2006). After
117 the reactivation, mitigation structures such as drainage systems and retaining walls, were built to
118 stabilize the landslide. Since then, no further deep-seated movements have occurred (Cervi et al.,
119 2012) until March 14th, 2016. During this reactivation, the flysch rocks belonging to the upper
120 part failed and deformed in a roto-translational movement and caused the failure of a retaining wall
121 (Fig. 2 b) and the mobilization of the landslide deposit in the lower part as a flow-like movement
122 (Fig. 2 c). The photos show that the deformation varied from several meters in the upper part
123 up to hundreds of meters in the lower part. The landslide mass slowed down towards the end of
124 March 2016.

125 In the middle of November 2017, it accelerated again: the upper earthflow deposits partially
126 reactivated and moved downslope. The first movements occurred in correspondence in the upper
127 part of the earthflow deposit with estimated displacements of several meters. The intensity of the
128 displacements gradually decreased in the lower portion. The velocity of the earthflow never reached
129 a value of zero because on February 20th, 2018 it was affected by another small acceleration and in
130 March 2018 changing geomorphological features like trenches, exposed material, and surface water
131 ponds demonstrated that it kept moving (Servizio Geologico Sismico e dei Suoli della Regione

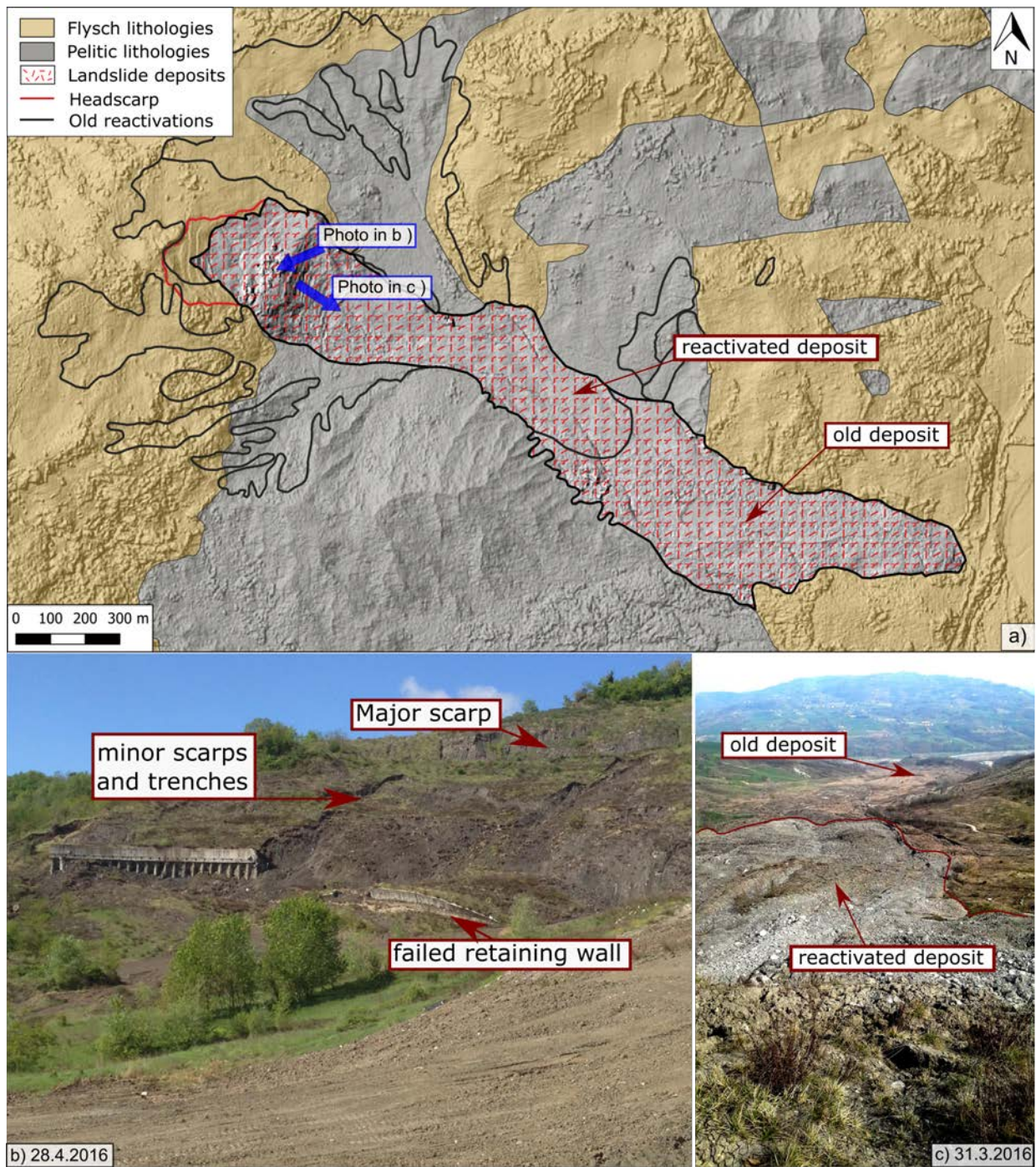


Figure 2: a) Map of the Ca Lita landslide with positions of the photos that were taken after the reactivation during March 2016. They show b) the rotational sliding in the upper part that caused the failure of the mitigation measurements (photo courtesy of Al Handwerger) and c) the flow like propagation in the lower part of the slope.

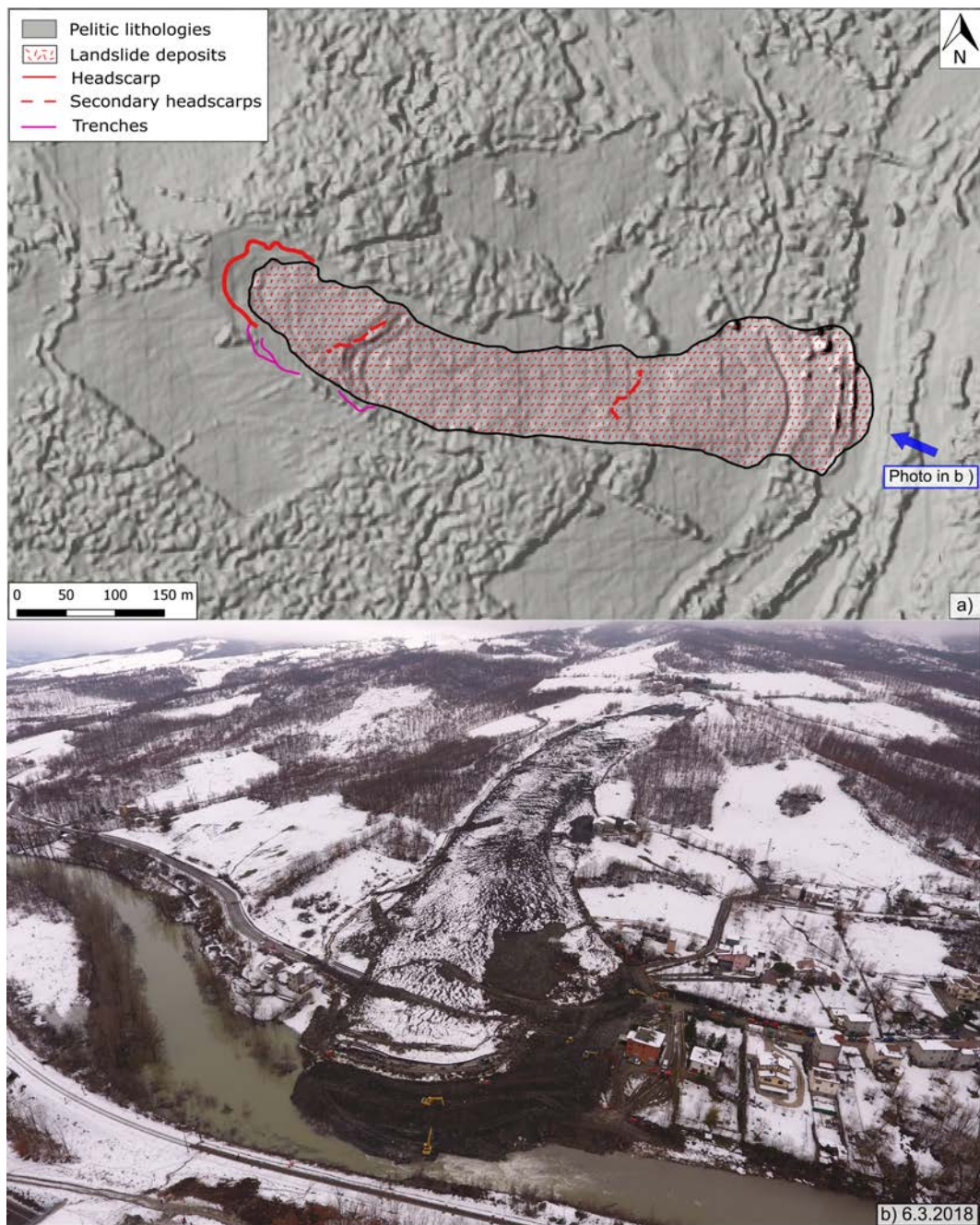


Figure 3: a) Map of the Marano earthflow with the deposits and the main morphological features related to the last reactivation being highlighted. b) Photo taken by a drone on the 6th of March 2018 (photo courtesy of Davide Marchioni).

133 2.2. *The Marano earthflow*

134 The Marano landslide (Fig. 1, 3) is located in the Bologna province (Italy) between 260 and
135 400 m a.s.l.; it is 700 m long and 100 m large for an estimated total volume of about 0.5 Mm^3 . The
136 landslide is a reactivated earthflow (Cruden and Varnes, 1996) that involve clay-shale lithologies
137 belonging to the Palombini Shale Formation (Panini et al., 2002). During the last century, it
138 reactivated twice: February 1996 and March 2018.

139 The 1996 event occurred after a period of intense rainfalls and snowfalls. The first motion
140 had been recorded on February 1st in the upper portion of the slope and rapidly propagated
141 downslope; after 6 days of rapid movement, it slowed down. The slope failure intersected different
142 infrastructures like roads, methane pipelines, phone and electricity lines (Servizio Geologico Sismico
143 e dei Suoli della Regione Emilia-Romagna, 2019). In the following period, mitigation strategies
144 were adopted including drainage systems in the earthflow deposits and gabions in the lower part
145 of the landslide to preserve the road below. For more than 20 years no signs of deformation
146 were observed on the rebuilt road, in the gabions or the earthflow area. On March 1st, 2018 the
147 landslide accelerated, however, after a period of snowmelt and rainfall. The mitigation structures
148 were destroyed and the deposits reached the Reno river which is well visible in Fig 3 b. The Marano
149 landslide moved with velocities of several meters per day for at least ten days, then decelerated. In
150 the following days, employees worked at the earthflow toe and removed a large amount of material
151 that was occluding the Reno river and also threatening the railways on the opposite bank.

152 3. Materials and Methods

153 3.1. *Synthetic aperture radar interferometry (InSAR): limitations and techniques*

154 Space-borne synthetic aperture radar interferometry (InSAR) is a remote sensing technique
155 that exploits the phase difference between two radar images that were acquired over a given track
156 of the earth surface by a satellite. Part of the phase difference is caused by the deformation of
157 the targets inside a pixel with respect to the sensor (Massonnet and Feigl, 1998; Rosen et al.,
158 2000; Bürgmann et al., 2000). It has been widely used for different applications in earth-sciences
159 including earthquakes (Fialko et al., 2005), land subsidence and uplift of aquifers (Schmidt and
160 Bürgmann, 2003; Chaussard et al., 2014) or because of glacial processes (Auriac et al., 2013, 2014),

161 volcanic deformation (Hooper et al., 2007), but also for landsliding (Bianchini et al., 2013; Handw-
162 erger et al., 2013; Raspini et al., 2019). InSAR, however, presents three major limitations which
163 are related to each other.

164

165 1. **Phase ambiguity:** The differential phase of an interferogram is ambiguous because it
166 is measured as a fraction of the wavelength and a deformation field will be mapped in
167 the range between $-\pi$ and π radians. At this stage, the interferometric phase is typically
168 called wrapped phase and in deforming areas a spatial pattern that is called interferometric
169 fringes can often be observed (Massonnet and Feigl, 1998, e.g.). The transgression from one
170 end of the spectrum to the other is occasionally also referred to as phase-jump. Resolv-
171 ing this phase-ambiguity to obtain absolute values requires a process that is called phase
172 unwrapping and in the past different techniques were proposed to address this problem
173 (e.g. Chen and Zebker, 2001; Hooper and Zebker, 2007). If the deformation field develops
174 over a small area, an under-sampling of the phase-jump may occur that will result in an
175 unwrapping error.

176 2. **Decorrelation:** One major draw-back of InSAR, especially in rural areas, is signal loss
177 that is also referred to as coherence loss or decorrelation of the interferogram (Zebker and
178 Villasenor, 1992). It occurs mainly when the surface between two acquisitions changes sig-
179 nificantly (temporal decorrelation), for instance because the timespan of the interferogram
180 is long, because deformation rates are high, because snow cover is present in one scene of
181 the interferogram or because vegetation starts to grow. Decorrelation may also occur if the
182 distance of the sensor between two acquisitions (known as perpendicular baseline) is large,
183 which is called baseline decorrelation. In the presence of noise from decorrelation also the
184 unwrapping will become more difficult and unwrapping errors will occur more frequently
185 (Chen and Zebker, 2001).

186 3. **Contributions to the interferometric phase:** Even if the interferometric phase is coher-
187 ent, it still contains contributions that can be regarded as noise if deformation is the main
188 goal of the analysis. The main sources of the undesired signal are the differential phase from
189 topography, atmosphere and orbital errors (Tarayre and Massonnet, 1996; Zebker et al., 1997;
190 Fattahi and Amelung, 2015).

191 Since the 1990s numerous space-borne SAR missions with three different wavelengths have
192 been active, ranging from short wavelength satellites with high spatial resolution of ca. 3 m
193 (3 cm X-BAND, COSMO-Skymed, TerraSAR-X) over C-Band with variable spatial resolutions
194 (5.6 cm, ERS, Envisat or Sentinel-1) to 23 cm L-Band sensors like JERS, ALOS PALSAR or ALOS
195 PALSAR-2 (see for instance Wasowski and Bovenga, 2014 for an exhaustive list). Because the
196 differential phase is measured as fraction of the satellite wavelength, short wavelength sensors are
197 potentially more sensitive to small displacements compared to long wavelength sensors, whereas
198 they will have more problems with decorrelation and unwrapping errors. Also, long wavelength In-
199 SAR datasets are known to maintain coherence well also in rural areas (e.g. Handwerger et al., 2013;
200 Schlogel et al., 2015), because longer wavelength radar waves can penetrate superficial vegetation
201 and even canopy (e.g. Prush and Lohman, 2014; Ni et al., 2014).

202 Different multitemporal techniques, like persistent scatterer interferometry (Ferretti et al., 2001;
203 Hooper et al., 2004), evolutions of it (Ferretti et al., 2011), small baseline techniques (Berardino
204 et al., 2002; Schmidt and Bürgmann, 2003) or hybrid approaches (Hooper, 2008), were developed
205 to address the problems of decorrelation and estimate different error terms of the phase. They
206 were frequently used in the past to infer spatiotemporal information of slope deformations both on
207 the scale of single slopes (Wasowski and Bovenga, 2014), as well as on larger scales (Raspini et al.,
208 2019). Near to the study area, small baseline techniques proved useful to assess tunneling induced
209 deformation (Bayer et al., 2017), but also the relationship between seasonal creep of landslides from
210 variations of precipitation (Bayer et al., 2018). All of the aforementioned works used techniques that
211 focused on extracting highly coherent pixels mostly on human structures, like houses or exposed
212 rock-outcrops. In the study area, however, most active landslides have a moderate vegetation cover,
213 rarely have exposed landslide material and only slow-moving deep-seated landslides have human
214 structures on them.

215 Similar geomorphological and geological conditions exist in Northern California, where only
216 long-wavelength data from ALOS permitted to reveal relationships between earthflow deformation
217 and the precipitation regime (Handwerger et al., 2013, 2015; Bennett et al., 2016), and in combi-
218 nation with offset tracking techniques also the slow down of earthflows because of extreme drought
219 conditions. Most recently, however, Handwerger et al. (2019) have shown that also the C-band
220 data acquired by Sentinel 1 can be successfully used to obtain high-quality interferograms on types

221 of landslide similar to the ones described in this paper.

222 *3.2. InSAR datasets and processing*

223 We performed interferometric processing of synthetic aperture radar images acquired by Coper-
224 nicus Sentinel 1 A/B satellites by using GMTSAR (Sandwell et al., 2011) and unwrapped the
225 complex interferograms with the Statistical-Cost, Network-Flow Algorithm (SNAPHU; Chen and
226 Zebker, 2001). The Sentinel images are C-band images (5.6 cm radar wavelength) acquired with
227 a minimum interval of acquisition of six days (12 days for each satellite, with a six days interval
228 between Sentinel 1A and Sentinel 1B). We studied the period between January 2015 and January
229 2019 by analyzing two descending orbits (south-moving satellites, looking west) and two ascending
230 orbits (north-moving satellites, looking east) for a total of four datasets for each landslide: Track
231 168, Track 95, Track 15 and Track 117 (Fig. 1). We initially processed a total of 869 interferograms
232 for the Marano landslide and 1419 interferograms for the Ca Lita landslide that were inspected
233 visually and only interferograms with a clear phase signal were considered for further processing.

234 The topographic phase was calculated and subtracted (e.g. Massonnet and Feigl, 1998; Bürgmann
235 et al., 2000) by using an external digital surface model (2x2 m DSM, provided by the Emilia Ro-
236 magna Region Services). Because of the small perpendicular baselines of Sentinel 1, the residual
237 DEM error is small compared to the signal from landslide motion and a correction scheme, like the
238 one proposed in (Fattahi and Amelung, 2015), proved not necessary. The large scale atmospheric
239 noise has been reduced by high pass filtering the interferograms and by selecting a stable reference
240 area close to the deforming region: we chose geomorphological (e.g. ridges) or anthropic features
241 (e.g. stable buildings) located near the landslides. Moreover, Gaussian and Goldstein filters (Gold-
242 stein and Werner, 1998) have been applied after the interferograms formation to reduce the noises
243 and enhance the deformation signal.

244 Despite the high acquisition frequency of Sentinel-1, unwrapping problems continued to arise
245 on the landslides during periods of high rates of displacement. Handwerger et al. (2015, 2019)
246 proposed a strategy to forward model the deformation to solve these unwrapping problems. We
247 adopted a similar approach that consisted in forming a deformation model by calculating the mean
248 rate of displacement from all interferograms without unwrapping errors. Then we used SNAPHU's
249 option that offers the possibility to subtract a deformation model before unwrapping adding it
250 back afterwards.

251 This approach helped to solve phase-jumps over the Marano landslide, whereas at the Ca Lita
252 landslide it only helped in few cases. This is probably because the Marano earthflow deformed
253 in a coherent slab, whereas the Ca Lita landslide has complex sliding features in the upper part
254 with high relative displacements and flow like deformation in the lower part with high absolute
255 displacements.

256 After this manual and iterative process of inspecting and improving interferograms, only those
257 without severe phase unwrapping problems were used to produce stacks of interferograms that
258 contain mean velocities and, in case of the Marano landslide, velocity time series.

259 We describe our results in terms of line-of-sight (LOS) displacement and velocity. Downslope
260 projection (Hilley et al., 2004) was not used to avoid the introduction of uncertainties deriving from
261 DEM-derived average slope and direction. In the case of our landslides, because of west-dipping,
262 moderately steep (10 to 20) slopes, positive and negative LOS displacements indicate downslope
263 movement for the ascending and descending orbit respectively. The downslope movement of earth-
264 flows is dominated by translation though vertical components can act at the toe or in the source
265 area (Picarelli et al., 2005).

266 4. Results

267 4.1. Spatial deformation patterns on the Ca Lita landslide

268 The kinematics of the Ca Lita earthflow are characterized by repeated variations of the rates
269 of displacement with values that exceed the detection limits of spaceborne radar interferometry.
270 Although at times it moves too fast to derive displacement or velocity time-series, a clear spatial
271 deformation pattern, roughly corresponding to the main landslide deposit, can be detected in a large
272 number of interferograms. Stacking series of interferograms, corresponding to a given time-interval,
273 increases the signal-to-noise ratio and highlights deforming features. The analysis and comparison
274 of successive interferometric stacks allow obtaining spatial and temporal information about the
275 landslide during phases of slow rates of displacement. During the failure stages, decorrelation and
276 unwrapping problems from fast-displacement can not be resolved, which is why the mean velocities
277 computed from the stacking process are locally underestimated. The spatial deformation signal,
278 however, is clear and can be used to document the evolution of the landslide movement just near
279 the activation stages.

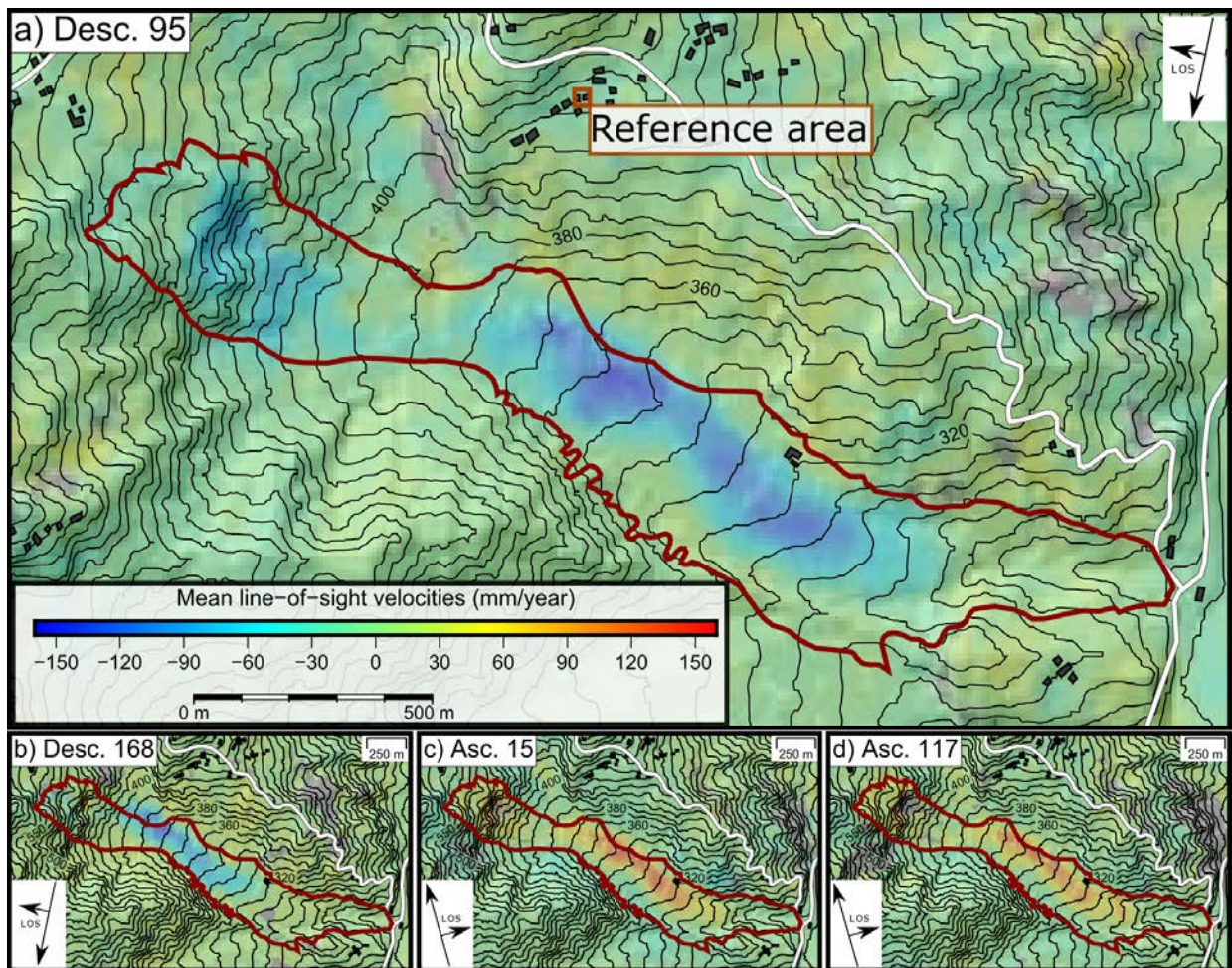


Figure 4: Ca Lita stacks of the entire interferograms series, concerning the period between January 2015 and January 2019: the numbers of the satellite Tracks are labeled at the top right of each image and the orbit directions are indicated at bottom right. Positive (red) values indicate motion away from the satellite along the line-of-sight and negative (blue) values indicate motion toward satellite along the line-of-sight. The numbers of interferograms that have been used to compute the stacks are: a) 286, b) 352, c) 372, d) 409. All the pixels averaging a value of coherence lower than 0.2 have been masked out in the figure and display grey color.

280 The stack of all manually-selected interferograms (from January 2015 to January 2019) high-
 281 lights the long-shaped morphology of the Ca Lita earthflow, that corresponds to the landslide
 282 deposit of the prior reactivation (Borgatti et al., 2006; Corsini et al., 2006; Servizio Geologico Sis-
 283 mico e dei Suoli della Regione Emilia-Romagna, 2019). The interferometric signal is particularly
 284 clear in the descending orbit 95 (Fig.4 a), whose stack indicates a range decrease and, hence, a
 285 movement towards the satellite. This observation is confirmed by descending orbit 168 (Fig.4 b).

286 Whereas the ascending orbits (117 and Fig.4 c and d) record range increases and, hence, movements
287 away from the satellite. In all cases, we used a mean coherence threshold of 0.25 to mask out areas
288 affected by low coherence. Because of the selection procedure of the interferograms, coherence is,
289 however, higher than 0.25 in most interferograms, which is why very few areas are masked out.

290 The difference between the ascending and the descending geometries should be interpreted as a
291 real deformation field that is oriented approximately down-slope. Maximum rates of displacement
292 are detected in the central part of the slope, where the type of movement transitions from sliding to
293 flowing. The landslide toe is relatively stable (no interaction with the national road was reported)
294 as well as the area above the crown, where houses are located, exhibit no deformation.

295 To document the temporal evolution of the Ca Lita landslide, we combined interferograms in
296 bimonthly stacks. We found that such frequency was suitable to resolve the different deformation
297 phases of this landslide. Fig. 8 reports the results derived from the descending orbit 168. The
298 failure of March 2016 is not clearly documented by radar interferometry because of persisting snow
299 cover in the area, which impeded to form coherent interferograms during this period. After failure,
300 the Ca Lita landslide exhibits enduring deformation: in the summer period the displacement signal
301 that is oriented towards the satellite is less evident and is located mainly the central portion of
302 the deposits (Fig.5 - a, b). In late fall of 2016 (Fig.5 b) and early 2017 (Fig.5 c) almost all the
303 landslide deposit is actively deforming.

304 At the beginning of 2017, the range of displacements decreases and are mainly located in the
305 central part of the slope where flow-like deformation is dominant and where the slope decreases
306 (Fig.5 c). In the upper part, small range increases were registered by the interferograms that span
307 this period. During the summer months, the slope was relatively stable with rates of LOS (Line-
308 Of-Sight) displacement lower than 100 mm/month. During September-October 2017 deformation
309 is intense (> 150 mm/year) and localized in the upper part of the landslide (Fig.5 d) where the
310 slope is relatively steep and sliding transitions into a flow-like type of movement. Following the
311 failure of November 2017, the whole landslide body, except for the toe, continued to move (Fig.5 e)
312 though rates of displacement appear generally lower. In the following period, the landslide activity
313 is clearly visible in the interferograms throughout the duration of our analysis. The LOS velocities
314 are locally sustained (> 150 mm/year), especially during the rainy season e.g., Nov.-Dec. 2018
315 stack in Fig.5 f).

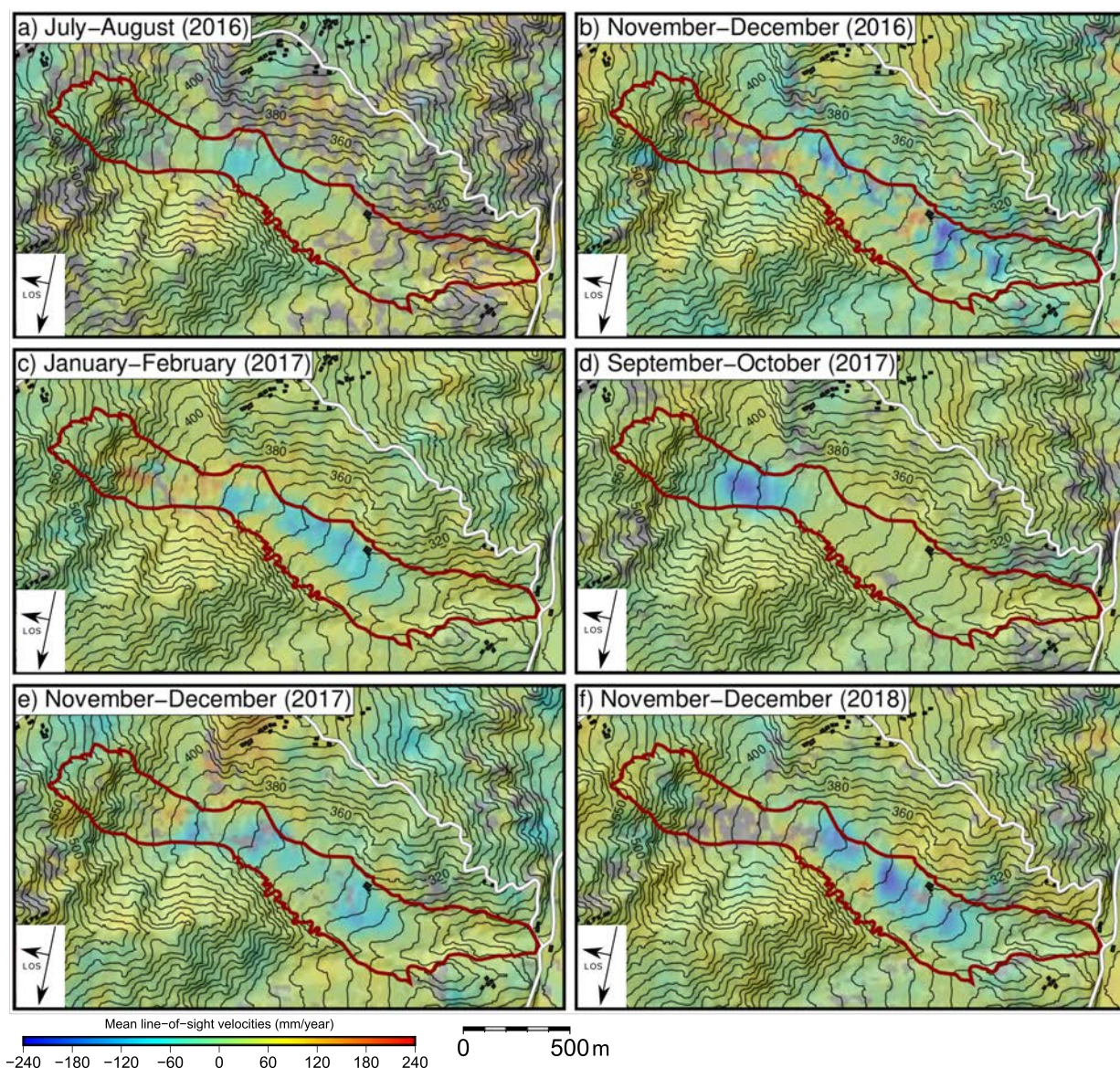


Figure 5: Two-months stacks for Track 168: a) July - August: 7 interferograms used for the stacking, b) November - December 2016: 10 interferograms used for the stacking, c) January - February 2017: 9 interferograms used for the stacking, d) September - October 2017: 8 interferograms used for the stacking, e) November - December 2017: 10 interferograms used for the stacking, f) November - December 2018: 10 interferograms used for the stacking. Three stages of movement can be observed: from a) to c) the deformation involves a very large portion of the deposits; in d) only the upper part in interested by displacements; from e) to f) the whole mass is involved again. Positive (red) values indicate motion away from the satellite along the line-of-sight and negative (blue) values indicate motion toward satellite along the line-of-sight. All the pixels averaging a value of coherence lower than 0.2 have been masked out in the figure and display grey color.

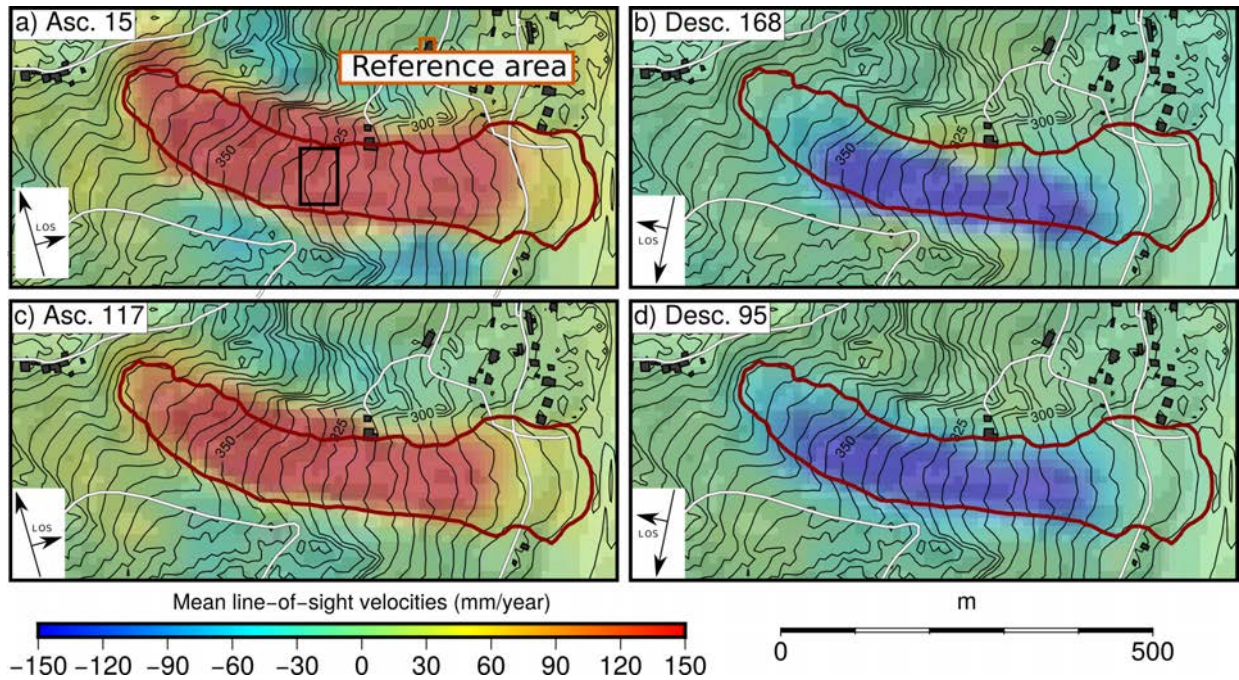


Figure 6: Marano stacks of the entire interferograms series, concerning the period between January 2015 and January 2019 derive from a) the ascending orbit 15 using 161 interferograms (the black box indicates the pixels that were used for the timeseries in figure 7, b) the descending orbit 168 using 218 interferograms c) the ascending orbit 117 using 209 interferograms and d) the descending orbit 95 with 281 interferograms. Warm colours indicate a movement away from the satellite along the line-of-sight, whereas cold colours indicate a movement towards the satellite. All the pixels averaging a value of coherence lower than 0.2 have been masked out.

317 The Marano earthflow reactivated catastrophically on March 1st, 2018 after 22 years of dor-
 318 mancy. The vast majority of selected interferograms (January 2015 to January 2019) detect active
 319 deformation along the slope. The apparently dormant landslide has been affected by detectable de-
 320 formation for at least two years before the catastrophic failure occurred. The stack of all manually-
 321 selected interferograms shows an extremely clear signal detected by all available orbits (Fig. 6).
 322 The reference area was chosen with respect to the houses of the locality Marano and the ascend-
 323 ing orbits show an almost identical spatial signal that indicates a range increase and, hence, a
 324 movement away from the satellite with more than 150 mm/year along the line-of-sight. The de-
 325 scending orbits, on the other hand, show a movement towards the satellite again with more than
 326 mm/year along the line-of-sight. This difference can be interpreted as a gravitational deformation
 327 oriented along the downslope direction.

328 The deformation signal shown by the interferometric stacks is consistent for all orbits (Fig. 6)
329 and indicates that surface displacements pervaded most of the landslide body with the exception
330 of the toe. This latter propagated downslope tens of meters during the paroxysmal phase, partially
331 damming the river and therefore being partially excavated. Our results, describe the pre- and
332 post-failure phase and therefore do not capture effects of rapid deformation. Also, the displacement
333 signal extending beyond the landslide perimeter in the ascending stacks (Fig. 6 a, 6 c) is compatible
334 with slow deformation of small slope portions that were not involved in the actual mapped failure.
335 Consider also that Sentinel spatial resolution is 20x5 m approximately and that noise-removal
336 spatial filtering further diminish the effective ground resolution.

337 Compared to the Ca Lita landslide, the interferometric signal on the Marano earthflow is less
338 noisy, because of lower rates of displacement , but also possibly because of the different kinematics.
339 Whereas the Ca Lita landslide is dominated by roto-translational sliding in the upper part and
340 flow-like deformation in the central and lower parts (Borgatti et al., 2006; Corsini et al., 2006), the
341 Marano landslide appears to move as a relative coherent block along slope-parallel slip surface/s.
342 Such response allowed us to successfully unwrap the Marano interferograms and extract velocity
343 information for the period between the beginning of Sentinel acquisition and the failure (Fig.7, a).

344 The velocity series are obtained by simply dividing the displacement of each interferogram by
345 the period between the two acquisitions that were used to form the interferogram. We used a
346 local regression analysis to fit the data and detect associated trends (line in Fig.7, a). Before the
347 launch of Sentinel 1B the frequency of velocity information is lower because only 12 and 24 days
348 interferograms are available and few are selected because of coherence issues. This is why the
349 trend before august 2016 is less defined. The most remarkable result is probably represented by
350 the regression lines of the four independent datasets that depict similar and coherent trends. To
351 interpret such trends, we compare them to the precipitation regime. The rainfall data have been
352 provided by the Regional Agency for Prevention, Environment and Energy of Emilia-Romagna
353 (Arpae) and the snowfalls data have been recorded at the Porretta station, respectively four and
354 eight kilometers far from the earthflow and at a comparable elevation. For each hydrological year
355 (starting in October) we calculated weekly rainfall values and cumulated precipitations (including
356 both rainfall and snowmelt; Fig.7 b).

357 During 2016 (October 2015 - September 2016) the only peak in velocity was resolved during

358 March, following a period of intense rainfalls: about 300 mm of rain occurred in the previous 60
359 days. In the following year intense snowmelt and rainfall cause the rates of LOS displacement to
360 exceed 100 mm/year in December 2016. In this case ascending datasets and descending dataset
361 95 capture the velocity peaking. During spring 2017 two peaks of high velocities were registered,
362 the first one occurred in March, while the second in May. The peaks are well registered by the
363 ascending dataset 117 and the descending dataset 95, whereas the other two orbits do retrieve
364 high rates of displacement during spring but do not resolve two distinct peaks. Again, the velocity
365 peaks follow two periods of precipitation with the first one being amplified by snowmelt.

366 During the dry summer of 2017, landslide velocities drop to almost null values along the line-of-
367 sight, but with the onset of hydrological year 2017-2018, the landslide acceleration started almost
368 synchronous with the first heavy precipitation of November 2017. The velocity continues to increase
369 systematically until the failure of March 2018. Both the peak velocities as well as the slope of the
370 velocity increases are higher compared to the previous years. Another difference between the
371 period that precedes the failure and the years 2015-2017 is the snowmelt significantly contributes
372 to an increase in the equivalent precipitation. The interferograms that directly precede or span
373 the failure are heavily decorrelated all over the Reno catchment because of the presence of snow
374 (3, b).

375 Once coherence is recovered (June 2018), the landslide is dormant and velocities are lower than
376 they were during the years 2015-2017.

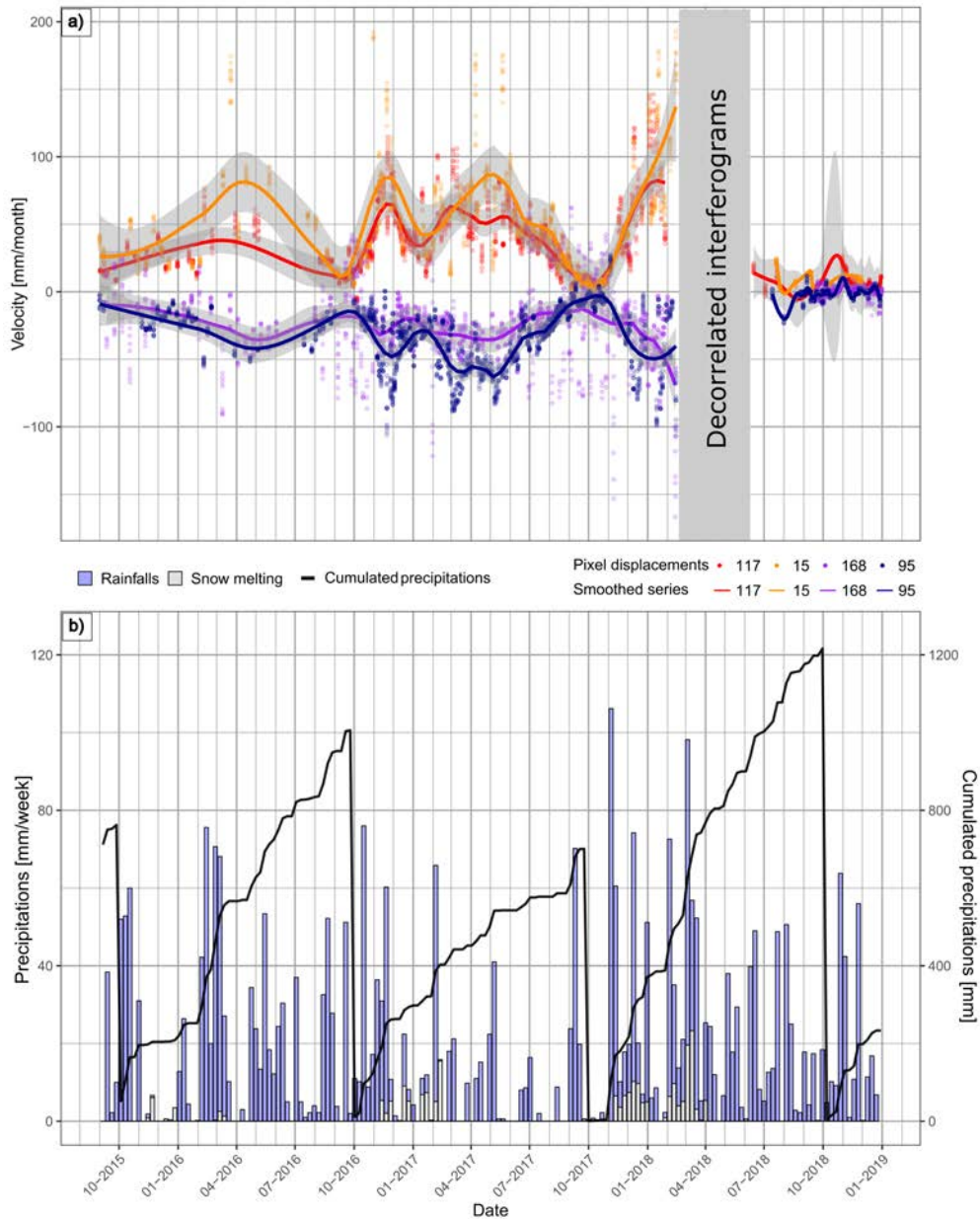


Figure 7: a) Velocity time series for each track at the Marano earthflow (positive values for ascending Tracks, negative values for descending Tracks). The dot symbols represent the pixels belonging to the investigated area (see Fig.6); the lines are derived by applying a local regression smoothing using the implementation of the ggplot package that takes into account a neighbourhood of 20 % w.r.t. to complete series. The gray bands are the 95 percent confidence interval of the smoothing operation. The gray box highlights the time period in which interferograms are completely decorrelated either because of the presence of snow or because of high rates of displacement during the failure. b) Weekly data of rainfall and snowmelt (left y-axes) are plotted together with the cumulated precipitation that contains rainfall and snow-melt(right y-axes): the series has been set to zero at the beginning of each hydrological year.

377 5. Discussion and conclusions

378 Interferometric analysis has been successfully applied to two slow-moving landslides that were
379 subject to generalized failures during the period of our investigations. Both landslides are char-
380 acterized by the scarce presence of man-made structures or rock-outcrops that could represent
381 stable scatterers in multi-temporal InSAR analysis. We used standard two-pass interferometry
382 (Handwerger et al., 2013, 2019) to detect deformation signals useful to document the evolution of
383 the landslides in the 2015-2019 interval. The InSAR data allow appreciating the spatial pattern of
384 deformation at successive time intervals.

385 In the case of the Ca Lita landslide, the deformation maps evidence inhomogeneous deforma-
386 tion fields throughout the landslide deposit that can be used to interpret the kinematics of the
387 phenomena. In fact, the pre-failure deformation at Ca Lita in 2017 was dominated by displace-
388 ments localized in the upper part of the slope. This is consistent with the dynamics described
389 for previous reactivation of this landslide (Borgatti et al., 2006; Corsini et al., 2006). Relatively
390 fast displacements are detected in space but obtaining quantitative results is associated with larger
391 uncertainties due to the presence of residual noise and unresolvable phase jumps. Though not
392 numerically accurate during the most active phases of landslide movement, interferograms, and
393 stacked interferograms contain useful information to: i) identify movement against surrounding
394 stable slopes; ii) document the spatial evolution of the movement.

395 At Ca Lita, different types of movement can be encountered (i.e. sliding in the upper part
396 and flow-like movement in the lower central and lower parts; Borgatti et al., 2006; Corsini et al.,
397 2006), displacement rates were often sustained in between the two failure episodes (March 2016
398 - December 2017) and possibly associated to high spatial small-scale variability due to flow-like
399 type of movement. Hence a velocity-time series similar to the one of the Marano case could not be
400 produced. A conceptual sketch in Fig. 8 a) illustrates that deformation exceeding approximately
401 120 mm/month cause signals in interferograms that are not correct from a numerical point of
402 view. This is because interferograms with higher displacement rates cause interferograms similar
403 to those in Fig. 8 b) that show multiple phase jumps in the landslide area. The signal can be
404 clearly attributed to deformation since coherence is high throughout the rest of the image. It is
405 however impossible to correctly count the interferometric fringes that occur in this interferogram
406 on the whole landslide body. There are also interferograms that are at the limit of decorrelation

407 and have one or two phase jumps (Fig. 8 c) which can be unwrapped by forward modeling the
408 deformation. Because the Ca Lita landslide has numerous open crevices and fissures along which
409 high differential displacements occurred, it is possible that unwrapping undersampled some phase
410 jumps.

411 From a geological point of view, the two analyzed landslide differ in several aspects. The
412 bedrock at Ca Lita is composed of flysch rocks in the upper part and chaotic clay shales in the
413 lower part of the slope, while the Marano earthflow is hosted only by chaotic clay shales. This
414 difference in the bedrock material might contribute to the fact that at the Ca Lita landslide different
415 kinematics coexist, while Marano is an earthflow like many others in the clay-shales rocks of the
416 Reno Catchment where flow-like morphology is associated to dominant sliding (Simoni et al., 2013).

417 The Marano earthflow remained in a dormant state for 20 years before it reactivated in March
418 2018. No damages were reported along the national road crossing the landslide at the toe nor by
419 the land owners upslope. However, InSAR data document active deformation for at least two years
420 before the failure occurred. Marano earthflow interferograms indicate the coherent displacement of
421 the landslide deposits. Velocity variations are well documented by the velocity time series that we
422 derive for each satellite track (Fig. 7). It has been possible to detect displacement rates ranging
423 from virtually null values to more than 100 mm/month. The velocity time series show repeated and
424 coherent velocity peaking that can be related to intense rainfalls and late summer velocity decline
425 observed during 2016 and 2017. The main triggering factor is the precipitation regime during
426 autumn 2017-spring 2018. The total amount of precipitation was significantly higher than average:
427 500 mm in the period between October 2017 and the failure (March 1st, 2018) most of them
428 (100 mm) in the 30 days preceding the failure (Fig. 7). Snow melting contributed to significantly
429 increase the equivalent precipitation during November and December 2017 and February 2018 when
430 we calculate 80 mm of snow melting that is added to 340 mm of rain. Also, the hydrological year
431 of 2017/2018 was preceded by an unusually dry summer which may have favored the formation
432 of fissures and cracks on the landslide body increasing permeability and hence the infiltration of
433 water (Malet et al., 2005).

434 From a technical point of view, results obtained on both cases show that standard InSAR can
435 deliver almost continuous deformation maps on landslides of the Northern Apennines that are
436 characterized by moderate vegetation and high displacement rates ranging from extremely slow to

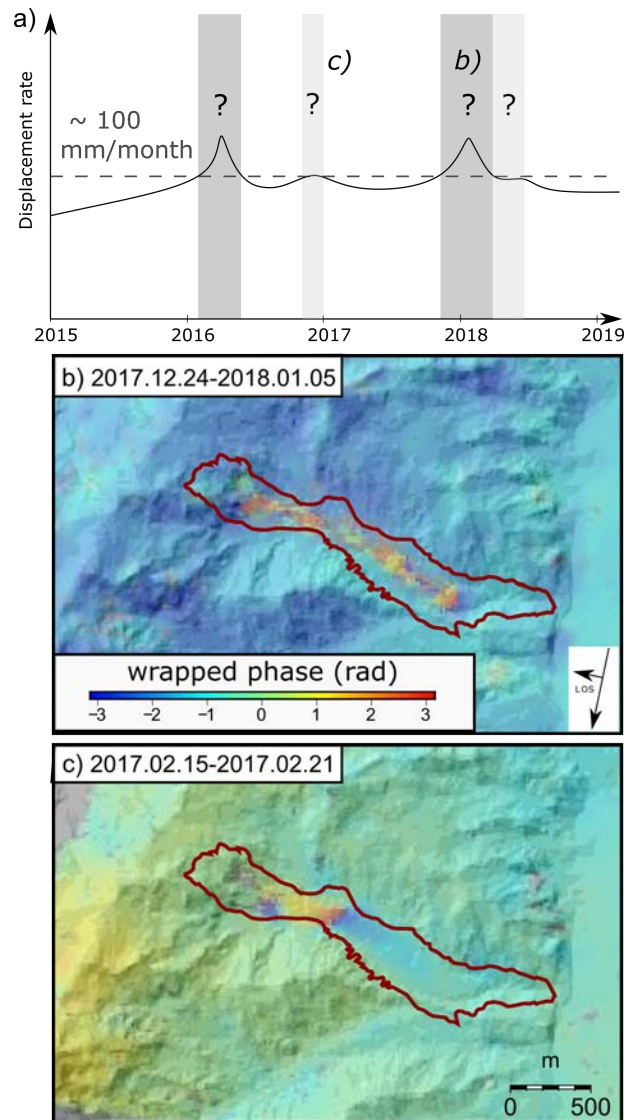


Figure 8: a) Conceptual sketch of the Ca Lita evolution. The dark gray boxes highlight periods in which several interferograms are decorrelated or display unwrapping problems because of fast displacement. The light gray boxes indicate periods in which displacement causes decorrelation in interferograms that span more than 12 days: the total number of decorrelated interferograms is slightly lower of that one referred to the dark gray boxes. The question marks in the gray boxes indicate the ambiguity of the the values for the rates of displacement if large decorrelation/unwrapping errors occur. In the other periods the velocity of the landslide is still often near to the upper limit and phase unwrapping may occur. b) Example of wrapped interferogram with multiple phase jumps close to complete decorrelation ("fast-displacement decorrelation"). c) Example of wrapped interferogram (Track 168) with only one phase jump that can be solved in the unwrapping step. The labels on the top right indicate the temporal baselines of the two examples.

437 about 100 mm/month. When the velocities approach the upper limit and/or the landslide shows
438 highly variable (pixel-scale) spatial deformation pattern, phase jumps cannot be further solved.
439 At lower values, velocities can be considered reliable though inherent uncertainties associated to
440 residual (topographic, atmospheric) noise remain. Despite the overall high quality of Sentinel 1
441 interferograms, we would like to remark the semi-quantitative significance of the displacement data
442 obtained from standard InSAR analysis. Residual noise due to topography and atmosphere can,
443 in fact, have a minor influence on the numerical displacement values that are obtained (Massonnet
444 and Feigl, 1998; Bürgmann et al., 2000). Bigger accuracy issues are caused by localized pixel-scale
445 shear zones resulting in phase jumps (Hu et al., 2019) and by displacement rates approaching the
446 limits of the technique (Rosen et al., 2000; Bürgmann et al., 2000).

447 The present work shows that InSAR-derived deformation maps supply a lot of information
448 about the spatial pattern and the temporal evolution of the landslides, also where stable reflectors
449 are scarce or absent. In our cases, the generalized failures of slow-moving, apparently dormant
450 earthflows were preceded by surface displacements that generated a clear interferometric signal.
451 On the contrary, such deformations did not cause evident damages and went undetected on the
452 ground. In a wider application perspective, we suggest that standard InSAR can provide qualitative
453 monitoring that can be used to detect and follow the evolution of landslide displacements preceding
454 (or following) a catastrophic failure. The high acquisition frequency of Sentinel 1 and the large
455 spatial extension of SAR scenes open also new perspectives in using this approach for large scale
456 analysis.

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